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Leptonic neutral-current probes in a short-distance DUNE-like setup

Based on 2402.00114 by SCCh, O. G. Miranda and J.W.F. Valle Phys. Rev. D 109, 115007

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Image: A math a math

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Neutrino	masses			

- Neutrino oscillations entering the precision (<1%) era. Simplest explanation for the L/E profile are massive neutrinos.
- For now, the neutrino mass mechanism is a mistery:
 - Tree level, radiative, extra dimensional origin...?
 - The scale(s) of relevant NP
 - Neutrino nature: Dirac or Majorana
 - Lepton number conservation/violation

• Neutrino mixing with new states could be the window to NP

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Image: A matrix

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- Let us assume that neutrinos mix with some new **heavy** gauge singlet states. This mixing can be naturally % level in low scale seesaw frameworks!
- Observational footprints are mainly
 - Collider direct production of m_{ν} mediators
 - cLFV (paradigmatically but not only $\mu^-
 ightarrow e^- \gamma)$
 - Non-standard effects in neutrino propagation
- None of these observations would be a statement on neutrino nature! See e.g. [1][2]

^[1]J. Bernabeu et al. "Lepton Flavor Nonconservation at High-Energies in a Superstring Inspired Standard Model". In: Phys. Lett. B 187 (1987), pp. 303–308. DOI: 10.1016/0370-2693(87)91100-2.

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Image: A matrix

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- 3 active neutrinos mix with *m* new (heavy) states
- The unitary mixing matrix is $(3 + m) \times (3 + m)$
- The upper 3 rows form a rectangular matrix K which characterizes the $\ell_{\alpha} W \nu_i$ interactions. See Concha's talk!

$$K = \begin{pmatrix} N & S \end{pmatrix}$$

$$K \, K^{\dagger} = 1_{3 imes 3}$$

Image: A matrix

• The Z boson interaction is characterized by the $(3 + m) \times (3 + m)$ matrix $P = K^{\dagger} K \neq 1$

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 If E ≪ M only the first 3 × 3 block of K and P can play a role, N and N[†]N, respectively.



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Important phenomenological consequences!

- CC can change flavour even at zero distance.
- NC is no longer diagonal.
- Observables at zero-distance depend on $(N^{\dagger}N)$. In the unitary limit this is the identity.
- Naive guess: Number of neutrino events in a given experiment is reduced compared to the unitary case. Not true!

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Basic formulation

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- N plays a central role in this setup
- We parametrize N as [3]

Non-unitarity formalism

$$N = \begin{pmatrix} \alpha_{11} & 0 & 0 \\ \alpha_{21} & \alpha_{22} & 0 \\ \alpha_{31} & \alpha_{31} & \alpha_{33} \end{pmatrix} \cdot U$$

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• In the unitary limit, N = I and U is the unitary matrix responsible for the standard oscillations.

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- Advantage of this parametrization: very clean theoretical interpretation in terms of mixing angles.
- The order of each parameter in the seesaw expansion is

$$\alpha_{ii}^2 \sim 1 - \varepsilon^2; \quad |\alpha_{ij}|^2 \sim \varepsilon^4$$

 [3]F. J. Escrihuela et al. "On the description of nonunitary neutrino mixing". In: Phys. Rev. D 92.5 (2015). [Erratum: Phys.Rev.D 93, 119905 (2016)], p. 053009. DOI: 10.1103/PhysRevD.92.053009. arXiv: 1503.08879 [hep-ph].

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Elastic scattering

Non-unitarity formalism

 We now consider the family of processes ν_α + e⁻ → ν_j + e⁻ and ν
_α + e⁻ → ν
_j + e⁻ at zero distance

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For concreteness, let's focus on the incoming muon neutrino case.



In the SM the cross section is given by

$$\left(\frac{d\sigma}{dT}\right)^{\mathsf{SM}} = \frac{2G_{\mu}^2m_e}{\pi} \left(g_L^2 + g_R^2\left(1 - \frac{T}{E\nu}\right)^2 - g_Lg_R\frac{m_eT}{E_{\nu}^2}\right)$$

And in the presence of non-unitarity

$$\left(\frac{d\sigma}{dT}\right)^{\mathsf{NU}} = \frac{\mathcal{P}_{\mu e}^{\mathsf{NC}}}{(NN^{\dagger})_{ee}(NN^{\dagger})_{\mu\mu}} \left(\frac{d\sigma}{dT}\right)^{\mathsf{SM}} + \frac{2m_e G_{\mu}^2}{\pi} \frac{\mathcal{R}e\left[\mathcal{P}_{\mu e}^{\mathsf{int}}\right]}{(NN^{\dagger})_{ee}(NN^{\dagger})_{\mu\mu}} \left\{\frac{\mathcal{P}_{\mu e}^{\mathsf{CC}}}{\mathcal{R}e\left[\mathcal{P}_{\mu e}^{\mathsf{int}}\right]} + 2g_L - g_R \frac{m_e T}{E_{\nu}^2}\right\}$$

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- The energy spectrum of the electron is modified!
- This feature is not shared by Cevens (pure NC) or inelastic scattering on nucleus (pure CC).
- However, this difference would be extremely hard to observe.
- It is also theoretically suppressed. Indeed, performing the the seesaw expansion and keeping only terms of $\mathcal{O}(\varepsilon)^2$ we find

$$\left(\frac{d\sigma}{dT}\right)^{\rm NU} \approx \left(2\alpha_{22}^2 - \alpha_{11}^2\right) \left(\frac{d\sigma}{dT}\right)^{\rm SM} + \mathcal{O}\left(\varepsilon^4\right)$$

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- If $E_{\nu}>$ 10 GeV we also have the purely CC process $\nu_{lpha}+e^-
 ightarrow
 u_j+\mu^-$
- For incoming μ neutrinos, the probability factor is given by

$$P_{\alpha\mu} = (N^{\dagger}N)_{ee}(N^{\dagger}N)_{\alpha\mu}(N^{\dagger}N)_{\mu\alpha}$$

$$\begin{aligned} \mathcal{P}_{\mu\mu} &\approx 2\alpha_{22}^2 + \alpha_{11}^2 - 2 \sim 1 - \mathcal{O}\left(\varepsilon^2\right) \\ \sigma &\approx \mathcal{P}_{\mu\mu}\frac{G_F^2}{\pi}\left(2\mathcal{E}_{\nu}m_e - m_{\mu}^2\right) \end{aligned}$$

Image: A matrix

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- $u_{\mu} + e
 ightarrow
 u_{j} + e$, mainly NC (at first order in seesaw expansion)
- $u_{\mu} + e \rightarrow \nu_j + \mu$, purely CC
- Final number of events could be bigger or smaller than the expected in the SM (due to the redefinition of G_F):

$$e^-$$
 events, NC: $rac{\#}{\#_{SM}} \approx 2lpha_{22}^2 - lpha_{11}^2 \sim 1 \pm \mathcal{O}(\varepsilon^2)$
 μ^- events, CC: $rac{\#}{\#_{SM}} \approx lpha_{22}^2 \sim 1 - \mathcal{O}(\varepsilon^2)$

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We will now analyze how the Near Detector (ND) of DUNE can constraint the NU parameters. However:

- Neutrino scattering with leptons has lower statistics compared to inelastic scattering with nucleus (way lower cross section)
- The experiment is not optimized to search for scattering on leptons
- Big flux uncertainties
- Constraints on NU from EW precision measurements will be stronger than those from neutrino physics constraints.

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- It is generally a good idea to find and explore complementary probes of a given phenomena
- Background under control (it is a cleaner process)
- Relatively less analyzed process compared to nucleus scattering

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Analysis				

- We compute the expected number of events in the SM for each flavour component of the flux (3.5 years per mode).
- We compute the NU (global) factors at order ε^2 .
- We extract the sensitivity on NU parameters.
- We compare them with current neutrino limits.



DUNE-like near detector, $\sigma = \sigma_{stat}$





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Take hon	ne ideas			

- Low scale seesaws: A well motivated and broad class of models leading to rich phenomenology. Non-unitarity effects can be at the % level.
- We have studied the effect of NU through the leptonic neutral current for the first time.
- The expected sensitivity will be competitive and complementary with other oscillation experiments (in particular on α²₁₁).

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Thanks!

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- A unitary matrix of order *n* can be parametrized with $\frac{n(n-1)}{2}$ angles and n(n+1) phases.
 - The complex rotation ω_{ij} has a mixing angle in the plane i j and a phase.
 - An (unphysical) diagonal matrix of phases times all the possible ω (in some order) parametrizes any unitary matrix

Parametrization of lphas in terms of mixing angles

- We can choose the ordering to be (NS \times NS)(S \times NS)($\omega_{23}\omega_{13}\omega_{12})$
- We identify the right hand side with the 'standard' mixing
- The first block cannot affect the 'active-active' 3×3 block
- The second block we subdivide in products of $\omega_{3j}\omega_{2j}\omega_{1j}$, which is lower tringular
- And a product of lower triangular matrices is also lower triangular
- diagonal entries are simply multiplications of cosines while off diagonal elements are proportional to sines (but are more complicated and include phases)

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• As a simple example in the 3 + 1 scheme we get

•
$$\alpha_{ii} = c_{i4}$$

•
$$\alpha_{ij} = s_{i4}s_{j4}e^{i(\phi_{i4}-\phi_{j4})}$$

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- $\nu_{\alpha} + N \rightarrow \nu_j + N$
- The COHERENT collaboration already detected the coherent scattering off nucleons in 2017 [4]
- Several experiments, including Coνus in Heidelberg [5]
- For future sensitivity analysis see e.g. [6]

^[4]D. Akimov et al. "Observation of Coherent Elastic Neutrino-Nucleus Scattering". In: Science 357.6356 (2017), pp. 1123-1126. DOI: 10.1126/science.aao0990. arXiv: 1708.01294 [nucl-ex].

^[5]H. Bonet et al. "Novel constraints on neutrino physics beyond the standard model from the CONUS experiment". In: JHEP 05 (2022), p. 085. DOI: 10.1007/JHEP05(2022)085. arXiv: 2110.02174 [hep-ph].

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Neutrino	limits			

- See the analysis in[7]
- Combines data from long (NOvA, T2K, MINOS) and short baseline (NOMAD, NuTeV) experiments
- At 90% CL:

$$\begin{aligned} 1 - \alpha_{11}^2 &\le 6 \times 10^{-2} \\ 1 - \alpha_{22}^2 &\le 1 \times 10^{-2}. \end{aligned}$$

[7]D. V. Forero et al. "Nonunitary neutrino mixing in short and long-baseline experiments". In: Phys. Rev. D 104.7 (2021), p. 075030. DOI: 10.1103/PhysRevD.104.075030. arXivi 2103.01998 [heg-ph] i きい 夏 - つくつ

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Number of events in the SM and probability factors

\mathcal{N}_{U}	$\nu \mod$	le	$\bar{\nu} \mod$	е	$\mathcal{N}_{\mathrm{NU}}/\mathcal{N}_{\mathrm{U}}$	Seesaw orde	r Main contribution
$\nu_a + e^- \rightarrow \nu_j + e^-$	\mathbf{events}	σ	\mathbf{events}	σ	${\cal P}G_F^2/G_\mu^2$		
$ u_e$	2.800	80	1.530	50	$2\alpha_{11}^2 - \alpha_{22}^2$	$1 \pm \mathcal{O}\left(\varepsilon^2\right)$	NC + CC
$ u_{\mu}$	31.400	700	5.800	100	$2\alpha_{22}^2 - \alpha_{11}^2$	$1 \pm \mathcal{O}\left(\varepsilon^2\right)$	NC
$ar{ u}_e$	430	20	780	30	$2\alpha_{11}^2 - \alpha_{22}^2$	$1 \pm O(\varepsilon^2)$	NC + CC
$ar u_\mu$	3.200	80	20.000 4	400	$2\alpha_{22}^2 - \alpha_{11}^2$	$1 \pm O\left(\varepsilon^2\right)$	NC
total	37.800	800	28.000	500			
\mathcal{N}_{U}	ν mo	de	$\bar{\nu}$ mo	de	$\mathcal{N}_{\mathrm{NU}}/\mathcal{N}_{\mathrm{U}}$	Seesaw order 1	Main contribution
$\mathcal{N}_{\mathrm{U}} \\ \nu_a + e^- \to \nu_j + \mu^-$	$\nu \mod \mathbf{events}$	de σ	$\bar{\nu} \mod \mathbf{events}$	de σ	${\cal N}_{ m NU}/{\cal N}_{ m U} onumber \ {\cal P}G_F^2/G_\mu^2$	Seesaw order 1	Main contribution
$\frac{\mathcal{N}_{\mathrm{U}}}{\nu_a + e^- \to \nu_j + \mu^-}$ $\frac{\nu_e}{\nu_e}$	$\nu \mod \frac{\nu}{0}$	de σ 0	$\bar{\nu} \mod \mathbf{events}$	de σ 0	$\frac{\mathcal{N}_{\rm NU}/\mathcal{N}_{\rm U}}{\mathcal{P} G_F^2/G_\mu^2}$ $ \alpha_{21} ^2$	Seesaw order 1 $\mathcal{O}\left(\varepsilon^{4}\right)$	$\frac{\text{Main contribution}}{\mathcal{O}\left(\varepsilon^{4}\right)}$
$ \frac{\mathcal{N}_{\mathrm{U}}}{\nu_{a} + e^{-} \rightarrow \nu_{j} + \mu^{-}} $ $ \frac{\nu_{e}}{\nu_{\mu}} $	ν mo events 0 17.900	de σ 0 400	 ν mo events 0 14.200 	de <i>σ</i> 	$\frac{\mathcal{N}_{\rm NU}/\mathcal{N}_{\rm U}}{\mathcal{P} G_F^2/G_\mu^2} \\ \frac{ \alpha_{21} ^2}{\alpha_{22}^2}$	Seesaw order $\begin{bmatrix} \mathcal{O}(\varepsilon^4) \\ 1 - \mathcal{O}(\varepsilon^2) \end{bmatrix}$	$ \begin{array}{c} \text{Main contribution} \\ \hline \mathcal{O}\left(\varepsilon^4\right) \\ \hline \text{CC} \end{array} $
$ \frac{\mathcal{N}_{\mathrm{U}}}{\nu_{a} + e^{-} \rightarrow \nu_{j} + \mu^{-}} $ $ \frac{\nu_{e}}{\nu_{\mu}} $ $ \bar{\nu}_{e}$	ν mo events 0 17.900 380	de σ 0 400 20	 ν mo events 0 14.200 230 	de σ 0 300 20	$\frac{\mathcal{N}_{\rm NU}/\mathcal{N}_{\rm U}}{\mathcal{P} G_F^2/G_\mu^2}$ $\frac{ \alpha_{21} ^2}{\alpha_{22}^2}$ $\frac{\alpha_{11}^2}{\alpha_{11}^2}$	Seesaw order 1 $\mathcal{O}(\varepsilon^4)$ $1 - \mathcal{O}(\varepsilon^2)$ $1 - \mathcal{O}(\varepsilon^2)$	$\begin{array}{c} \text{Main contribution} \\ \hline \\ \mathcal{O}\left(\varepsilon^{4}\right) \\ \\ \text{CC} \\ \\ \text{CC} \\ \\ \end{array}$
$ \begin{array}{c} \mathcal{N}_{\mathrm{U}} \\ \nu_{a} + e^{-} \rightarrow \nu_{j} + \mu^{-} \\ \hline \nu_{e} \\ \bar{\nu}_{e} \\ \bar{\nu}_{\mu} \\ \hline \bar{\nu}_{\mu} \end{array} $	ν mo events 0 17.900 380 0	de σ 10 400 20 0	$\bar{\nu} \mod \frac{\bar{\nu}}{230}$	de σ 300 20 0	$\frac{\mathcal{N}_{\rm NU}/\mathcal{N}_{\rm U}}{\mathcal{P} G_F^2/G_\mu^2} \\ \frac{ \alpha_{21} ^2}{\alpha_{22}^2} \\ \alpha_{11}^2 \\ \alpha_{21} ^2 \\ \end{array}$	Seesaw order 1 $\mathcal{O}(\varepsilon^4)$ $1 - \mathcal{O}(\varepsilon^2)$ $1 - \mathcal{O}(\varepsilon^2)$ $\mathcal{O}(\varepsilon^4)$	$\begin{array}{c} \text{Main contribution} \\ \hline \mathcal{O}\left(\varepsilon^{4}\right) \\ \hline \text{CC} \\ \text{CC} \\ \text{CC} \\ \mathcal{O}\left(\varepsilon^{4}\right) \end{array}$

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Comparison with existing constraints

- Oscillations
 - Long baseline experiments
 - Short baseline (zero distance)
- Lepton flavour universality
 - π & K decays into μ^- and e^-
 - τ^- decays (hadrons or leptons) \leftarrow function of (α_{33})
 - eta decays and CKM unitarity 🎇
- EW precision observables
 - *W* mass, *s_W*, Γ_Z...
 - CDF-II W mass [8]

[8]Mattias Blennow et al. "Right-handed neutrinos and the CDF || anomaly". In: Phys. Rev. D 106.7 (2022), p. 073005. DOI: 10.1103/PhysRevD.106.073005. arXiv: 2204.04559 [hep@ph]. < ≧ → < ≧ → < ≧ → < <

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LFU				

- See [9] for a nice update
- We can translate the decay rates of pions (or Kaons) to electrons and muons into couplings with the W. The experimental result

$$\left(\frac{g_e}{g_{\mu}}\right)^2 = 0.998 \pm 0.002$$

• In the SM this ratio is 1. In the presence of non-unitarity

$$\left(\frac{g_e}{g_{\mu}}\right)^2 = 1 + \alpha_{11}^2 - \alpha_{22}^2$$

 [9]Douglas Bryman et al. "Testing Lepton Flavor Universality with Pion, Kaon, Tau, and Beta Decays".

 In: Ann. Rev. Nucl. Part. Sci. 72 (2022), pp. 69-91. DOI: 10.1146/annur.ev-nucl-110121-051223. arXiv:

 2111.05338 [hep-ph].

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• We can also compare the effective coupling of β and μ decays. In the SM

$$\left(rac{G_eta}{G_\mu}
ight)^2 = \sum_i |V_{ui}|^2 = 1$$

• Different measurements of nuclear processes give

$$\sum_{i} |V_{ui}|^2 = 1 - (19.5 \pm 5.3) \times 10^{-4}$$

• Known as the **Cabibbo anomaly**. This anomaly only gets worse in the presence of (leptonic) non-unitarity

$$\left(\frac{\textit{G}_{\beta}}{\textit{G}_{\mu}}\right)^2 = 2 - \alpha_{22}^2 > 1$$

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 Including all the LFU and EW precision measurements gives a much stronger constraints than the ones obtained from DUNE-PRISM or oscillations [10]

> 95% CL: $1 - \alpha_{11}^2 \le 2 \times 10^{-3}$ $1 - \alpha_{22}^2 \le 2 \times 10^{-4}$

 Caveat: The constraints are pushed towards zero due to the Cabibbo anomaly. Would be interested in seeing a similar analysis excluding the CKM unitarity test.

^[10] Iattias Blennow et al. "Bounds on lepton non-unitarity and heavy neutrino mixing". In: JHEP 08 (2023), p. 030. DOI: 10.1007/JHEP08(2023)030. arXiv: 2306.01040 [hep-ph]. « @ » « ≧ » « ≧ » ≧ → Q @ Salvador Centelles Chuliá Erice 2024, Italy

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Background reduction

We can take advantage of the fact that the electron scattering will be mainly forward. See for example[11]



[11]. Zazueta et al. "Improved constraint on the MINER ν A medium energy neutrino flux using $\nu^-e^- \rightarrow \nu^-e^-$ data". In: *Phys. Rev. D* 107.1 (2023), p. 012001. DOI: 10.1103/PhysRevD.107.012001. arXiv: 2209.05540 [hep-ex].

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Redefinition of G_F

Non-unitarity formalism

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- G_F is measured through the μ^- decay.
- Lepton non-unitarity modifies this decay. The effective muon decay G_{μ} is related to the 'real' G_F by

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• Typically suppressed by the 'active-heavy' mixing.

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High and low scale seesaws

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• The type-I seesaw is a paradigmatic example of a high scale seesaw



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Image: A matrix



$$M_{\nu} = \begin{pmatrix} 0 & Y v \\ Y^{T} v & M \end{pmatrix}, \ U^{T} M_{\nu} U = m_{d}, \ U = \begin{pmatrix} N & S \\ X & Y \end{pmatrix}$$

- N and S are the 'active-active' and 'active-sterile' mixings. By construction $N^{\dagger}N = 1 S^{\dagger}S$.
- In the seesaw expansion we loosely define the parameter $\varepsilon \sim O(Y v/M)$ and diagonalize perturbatively, finding

$$m_{\nu} = v^2 Y^T M^{-1} Y, \quad S^* = Y v M^{-1} V \sim \varepsilon, \quad \varepsilon^2 \sim O(m\nu/M)$$

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• The inverse seesaw is a paradigmatic example of a low scale seesaw.



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Inverse seesaw				

• Introduces a new scale $\mu \ll \Lambda_{\rm EW}$

$$M_{\nu} = \begin{pmatrix} 0 & Y v & 0 \\ Y^{T} v & \mu' & M \\ 0 & M^{T} & \mu \end{pmatrix}$$

• In the seesaw expansion (one gen)

$$m_{\nu} = rac{Y^2 v^2}{M^2} \mu, \quad S = \begin{pmatrix} rac{Y v}{M^2} \mu & rac{Y v}{M} \end{pmatrix} \sim \begin{pmatrix} rac{m_{
u}}{Y v} & arepsilon \end{pmatrix}$$

• The second component can be % level, even if $m_
u o 0$

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Image: A mathematical states and a mathem

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α parametrization Neutrino limits SM prediction Non-neutrino constraints Background Indirect effects

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The DUNE experiment



- DUNE (Deep Underground Neutrino Experiment) is an ambitious neutrino experiment under construction. Will determine the mass ordering and improve precision on θ_{23} , δ_{CP} and θ_{13} [12]
- Two beam modes (neutrino/antineutrino), mainly u_{μ} or $ar{
 u}_{\mu}$

[12]. Abi et al. "Long-baseline neutrino oscillation physics potential of the DUNE experiment". In: Eur. Phys. J. C 80.10 (2020), p. 978. DOI: 10.1140/epjc/s10052-020-08456-四. arXi傍 2006.36043 [hep-ex] のへつ

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The near detector

- Mainly for cross-checking the neutrino flux, but we can use it to do BSM analysis too
- Will be the first purely leptonic test of "zero distance neutral oscillations"

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