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based on works: [1] PRL 129(15): 151601, (2022) [2] PLB 840, 137839, (2023) [3] PRD 108, 12, L121701 (2023) [4] PRD 109, 10, 105001 (2024)

Hydrodynamic Manifestations of GRAVITATIONAL CHIRAL Anomaly

# CONTENTS

- Introduction
- General derivation
- Direct verification: spin 1/2
- Direct verification: spin 3/2 (RSA model)
- Conclusion

# PART 1 INTRODUCTION

# Gravity in flat spacetime: Cheshire cat GRIN

**"Well! I've often seen a cat without a grin,' thought Alice 'but a grin without a cat! It's the most curious thing i ever saw in my life!"**

― Lewis Carroll, Alice in Wonderland



# Gravity in flat spacetime: Cheshire cat GRIN

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#### CVE and CME - new anomalous transport



Derivation **without entropy** current and generalization to the **second order** in gradients:

**[Shi-Zheng Yang, Jian-Hua Gao, Zuo-Tang Liang, Symmetry 14 (2022) 5, 948]**

**[M. Buzzegoli, Lect. Notes Phys. 987, 53–93 (2021)]** Use **global equilibrium**

#### Modern development and the problem

What about the **gravitational chiral anomaly**?

• The gravitational chiral anomaly (unlike gauge part) grows **rapidly** with **spin**:

$$
\langle \nabla_{\mu} \hat{j}_{A}^{\mu} \rangle_{S} = \frac{(S - 2S^{3})}{96\pi^{2} \sqrt{-g}} \varepsilon^{\mu\nu\rho\sigma} R_{\mu\nu\kappa\lambda} R_{\rho\sigma}^{\quad \kappa\lambda}
$$

**[M. J. Duff, in First School on Supergravity (1982) arXiv:1201.0386]**

**[S. M. Christensen, M. J. Duff, Nucl. Phys. B 154, 301–342 (1979)]**

- How does the **gravitationa**l chiral anomaly manifest itself in **hydrodynamics**?
- Is it possible to see the **cubic factor**  $S-2S^3$  in hydrodynamics?

# GENERAL DERIVATION

#### Decomposition of the tensors

● **Components of the thermal vorticity tensor**

**6** components

**[M. Buzzegoli, E. Grossi, F. Becattini, JHEP 10 (2017) 091]**

$$
\varpi_{\mu\nu}=\epsilon_{\mu\nu\alpha\beta}w^\alpha u^\beta+\alpha_\mu u_\nu-\alpha_\nu u_\mu
$$

Similar to the expansion for the electromagnetic field

● We also decompose the **Riemann tensor** into the **components:**

$$
R_{\mu\nu\alpha\beta} = u_{\mu}u_{\alpha}A_{\nu\beta} + u_{\nu}u_{\beta}A_{\mu\alpha} - u_{\nu}u_{\alpha}A_{\mu\beta} - u_{\mu}u_{\beta}A_{\nu\alpha}
$$
  
\n
$$
+ \epsilon_{\mu\nu\lambda\rho}u^{\rho}(u_{\alpha}B^{\lambda}{}_{\beta} - u_{\beta}B^{\lambda}{}_{\alpha})
$$
  
\n20 components  
\n
$$
+ \epsilon_{\alpha\beta\lambda\rho}u^{\rho}(u_{\mu}B^{\lambda}{}_{\nu} - u_{\nu}B^{\lambda}{}_{\mu})
$$
  
\n
$$
+ \epsilon_{\mu\nu\lambda\rho}\epsilon_{\alpha\beta\eta\sigma}u^{\rho}u^{\sigma}C^{\lambda\eta}
$$

Coincide with **3d** tensors in the fluid rest frame:

**[L. D. Landau and E. M. Lifschits, The Classical Theory of Fields, Vol. 2, 1975] [A. Z. Petrov, 1950]**

• We consider **Ricci-flat** spaces  $R_{\mu\nu}=0$ 

#### Gradient expansion in the curved spacetime

The **gravitational chiral anomaly** has the **4th order** in gradients – it is to be related to the **3rd order** terms in gradient expansion of the axial **current**.

Hydrodynamic expansion of the axial current up to the **3rd order** in gradients:



See also gradient expansion for the fluid in the gravitational field, e.g.:

**[P. Romatschke, Class. Quant. Grav. 27, 025006 (2010)]**

**[S. M. Diles, L. A. H. Mamani, A. S. Miranda, V. T. Zanchin, JHEP 2020, 1–40 (2020)]**

#### Anomaly matching: principle

Following **[D.T. Son, P. Surowka, PRL, 103 (2009) 191601]**

- it would be necessary to construct the **entropy current**.

However in **[Shi-Zheng Yang, Jian-Hua Gao, and Zuo-Tang Liang, Symmetry 14, 948 (2022)]**

it is shown that it is possible to use the **global equilibrium** condition

 $\nabla_{\mu}\beta_{\nu}+\nabla_{\nu}\beta_{\mu}=0$ 

After that it is enough to consider **only** the equation for the current.

● **Good for gravity,** which is **complicated** in general case!

**We use only:**

$$
\nabla_{\mu}j^{\mu}_A=\mathscr{N}\epsilon^{\mu\nu\alpha\beta}R_{\mu\nu\lambda\rho}{R_{\alpha\beta}}^{\lambda\rho} \textbf{.}
$$

**Substitute** the gradient expansion:

 $\nabla^{\mu} (\xi_1(T) w^2 w_{\mu} + \xi_2(T) \alpha^2 w_{\mu} + \xi_3(T) (\alpha w) w_{\mu} + \xi_4(T) A_{\mu\nu} w^{\nu} + \xi_5(T) B_{\mu\nu} \alpha^{\nu}) = 32 \mathcal{N} A_{\mu\nu} B^{\mu\nu}$ 

### Anomaly matching: system of equations

This **system** of linear **differential** equations has the form:

$$
-3T\xi_1 + T^2\xi_1' + 2T\xi_3 = 0
$$
  
\n
$$
-3T\xi_2 + T^2\xi_2' - T\xi_3 + T^2\xi_3' = 0
$$
  
\n
$$
T^2\xi_4' + 3T\xi_5 + 2T^{-1}\xi_2 + T^{-1}\xi_3 = 0
$$
  
\n
$$
-2T^{-1}\xi_1 - 3T\xi_4 - T\xi_5 = 0
$$
  
\n
$$
T^2\xi_5' - T\xi_5 - T^{-1}\xi_3 = 0
$$
  
\n
$$
-T^{-1}\xi_4 + T^{-1}\xi_5 - 32\mathcal{N} = 0
$$
  
\n
$$
\downarrow
$$
<

#### **SOLUTION**

Let's move on to the limit of **flat space-time**. Despite the absence of a gravitational field, there **remains** a contribution to the axial current induced by the **gravitational chiral anomaly:**



● A new type of anomalous transport – the **Kinematical Vortical Effect (KVE)**. Does not explicitly depend on temperature and density → determined only by the **kinematics** of the flow.

# DIRECT VERIFICATION: spin 1/2

# Transport coefficients and anomaly: spin 1/2

● **KVE** In **[GP, O.V. Teryaev, and V.I. Zakharov, JHEP, 02:146, 2019], [V. E. Ambrus, JHEP, 08:016, 2020], [A. Palermo, et al. JHEP 10 (2021) 077]** and for  $\omega^3$  in **[A. Vilenkin, Phys. Rev., D20:1807–1812, 1979]** the following expression was obtained:

● Comparing it with the well-known anomaly **[L. Alvarez-Gaume, E. Witten, Nucl. Phys., B234:269, 1984]**:

$$
j_{\mu}^{A} = \left(\frac{T^{2}}{6} + \frac{\mu^{2}}{2\pi^{2}} - \frac{\omega^{2}}{24\pi^{2}} - \frac{a^{2}}{8\pi^{2}}\right)\omega_{\mu}
$$

$$
\nabla_{\mu} j^{\mu}_{A} = \frac{1}{384\pi^2 \sqrt{-g}} \varepsilon^{\mu\nu\alpha\beta} R_{\mu\nu\lambda\rho} R_{\alpha\beta}{}^{\lambda\rho}
$$

We see that the formula is **fulfilled**:

$$
\frac{\lambda_1 - \lambda_2}{32} = \mathcal{N}
$$
 
$$
\longrightarrow \left( -\frac{1}{24\pi^2} + \frac{1}{8\pi^2} \right) / 32 = \frac{1}{384\pi^2}
$$

#### **Correspondence** between **gravity** and **hydrodynamics** is **confirmed!**

# DIRECT VERIFICATION: SPIN 3/2

## Rarita-Schwinger-Adler model of spin 3/2

The **Rarita-Schwinger theory** – well-known theory of spin 3/2. But this theory has a number of **pathologies.**

● For example, it **doesn't allow to construct perturbation theory!**

Solved in **[Stephen L. Adler. Phys. Rev. D, 97(4):045014, 2018]** by introducing of interaction with additional spin  $\frac{1}{2}$  field:



**Anomaly** was found in **[Prokhorov, Teryaev, Zakharov, Phys.Rev.D 106 (2022) 2, 025022]**

$$
\nabla_{\mu} j^{\mu}_A = -\frac{19}{384\pi^2 \sqrt{-g}} \varepsilon^{\mu\nu\rho\sigma} R_{\mu\nu\kappa\lambda} R_{\rho\sigma}{}^{\kappa\lambda}
$$

**–19 times** different from the anomaly for spin  $\frac{1}{2}$ 

### Zubarev density operator

Global Equilibrium Conditions

$\nabla_{\mu}\beta_{\nu} + \nabla_{\nu}\beta_{\mu} = 0$	$\nabla_{\mu}\zeta = 0$	Thermal vorticity tensor
Form of the density operator for a medium with rotation and acceleration	Ans. $\widehat{\rho} = \frac{1}{Z} \exp \left[ -b_{\mu}\widehat{P}^{\mu} + \frac{1}{2}\varpi_{\mu\nu}\widehat{J}^{\mu\nu} + \zeta \widehat{Q} \right]$	Lorentz Transform
$\widehat{\omega}_{\mu\nu}\widehat{J}^{\mu\nu} = -2\alpha^{\rho}\widehat{K}_{\rho} - 2w^{\rho}\widehat{J}_{\rho}$	Lorentz Transform	
$\widehat{K}^{\mu}$ - boost (related to acceleration)	$\widehat{J}^{\mu\nu} = \int_{\Sigma} d\Sigma_{\lambda} (x^{\mu}\widehat{T}^{\lambda\nu} - x^{\nu}\widehat{T}^{\lambda\mu})$	

### KVE in RSA theory: calculation

• Our goal is to calculate the conductivities  $\lambda_1$  and  $\lambda_2$  in the KVE current:

$$
j_{A,KVE}^{\mu} = \lambda_1(\omega_{\nu}\omega^{\nu})\omega^{\mu} + \lambda_2(a_{\nu}a^{\nu})\omega^{\mu}
$$

 $+1)\pi T$ 

• Using the perturbation theory, we obtain:

$$
\lambda_1 = -\frac{1}{6} \int_0^{|\beta|} [d\tau] \langle T_{\tau} \hat{J}_{-i\tau_x}^3 \hat{J}_{-i\tau_y}^3 \hat{J}_{-i\tau_z}^3 \hat{j}_A^3(0) \rangle_{T,c}
$$

● The matrix element has the form of a product of **vertices** and **propagators**.

#### **Vertices**

$$
\mathcal{D}_{(ij)}^{\mu\nu} = -\frac{1}{2}(-i)^{\delta_{0\mu} + \delta_{0\nu}} \varepsilon^{ij\nu\beta} \left( \gamma_5 \tilde{\gamma}_{\mu} \tilde{\partial}_{\beta}^{X_2} - \frac{1}{4} \gamma_5 \tilde{\gamma}_{\beta} [\tilde{\gamma}_{\vartheta}, \tilde{\gamma}_{\mu}] \left( \tilde{\partial}_{\vartheta}^{X_1} + \tilde{\partial}_{\vartheta}^{X_2} \right) \right) + (\mu \leftrightarrow \nu)
$$
  
0 \le (i, j) < 4

#### **Propagators**

$$
\langle T_{\tau}\tilde{\psi}_{a\mu}(X_{1})\tilde{\bar{\psi}}_{b\nu}(X_{2})\rangle_{T} = \sum_{P} e^{iP_{\alpha}(X_{1}-X_{2})^{\alpha}} \frac{i}{2P^{2}} \left(\tilde{\gamma}_{\nu}\tilde{P}\tilde{\gamma}_{\mu} + 2\left[\frac{1}{m^{2}} - \frac{2}{P^{2}}\right]P_{\mu}P_{\nu}\tilde{P}\right)_{ab}
$$
  

$$
\langle T_{\tau}\tilde{\psi}_{a\mu}(X_{1})\bar{\lambda}_{b}(X_{2})\rangle_{T} = \sum_{P} e^{iP_{\alpha}(X_{1}-X_{2})^{\alpha}} \frac{-P_{\mu}\tilde{P}_{ab}}{mP^{2}}
$$
  
Mixed terms are non-zero  
here  $P_{\mu} = (p_{n}, -\mathbf{p}), p_{n} = (2n$   
Field  $\lambda$  is non-propagating

## KVE vs Gravitational Anomaly



- The relationship between the transport coefficients in a **vortical accelerated fluid** and the **gravitational** chiral **anomaly** is shown!
- Verification in a **nontrivial** case with higher spins and interaction.

# CONCLUSION

### **CONCLUSION**

- The relationship between the hydrodynamic current in the third order of gradient expansion  $\lambda_1(\omega_\nu \omega^\nu) \omega_\mu$  and  $\lambda_2(a_\nu a^\nu) w_\mu$ , the Kinematic Vortical Effect (KVE), and the gravitational chiral anomaly has been established.
- The obtained formula has been verified directly for spins **1/2** and **3/2**.

# Additional

# slides

# Generalization to (anti)de Sitter space

• Going **beyond approximation**  $R_{\mu\nu} = 0$ **[Khakimov, Prokhorov, Teryaev, Zakharov,2401.09247 (2024)]**

 $\frac{\lambda_1 - \lambda_2}{32}$  = N -- anomaly-hydro relation remains valid

 $j_\mu^A \sim a^2 \omega_\mu \ll \frac{1}{2}$   $j_\mu^A \sim R \omega_\mu$  -- equivalence principle in higher orders

● **5-dimensional Unruh effect**:

**[Khakimov, Prokhorov, Teryaev, Zakharov, Phys.Rev.D 108, 12, L121701 (2023)]**

The temperature measured by an accelerated observer in **(A)dS space** is determined by the 5-dimensional acceleration!

#### **[S. Deser and O. Levin, Phys. Rev. D, 59:064004, 1999]**

Hydrodynamic expansion for the stress-energy tensor:



$$
\langle \hat{T}^{\mu\nu} \rangle_{s=1/2} = \left[ \frac{7\pi^2}{180} T^4 + \frac{1}{72} \left( |a|^2 + \frac{R}{12} \right) T^2 - \frac{17}{2880\pi^2} \left( |a|^2 + \frac{R}{12} \right)^2 \right] \left( 4u^\mu u^\nu - g^{\mu\nu} \right) + \frac{11}{960\pi^2} \left( \frac{R}{12} \right)^2 g^{\mu\nu}
$$

 $\langle \hat{T}^{\mu\nu}\rangle (T=T_{UR})=\frac{k}{4}R^2g^{\mu\nu}$  has a vacuum form  $T_{UR} = \frac{a_5}{2\pi} = \sqrt{T_U^2 + T_R^2} = \frac{\sqrt{|a|^2 + R/12}}{2\pi}$ 

### Rarita-Schwinger-Adler model of spin 3/2

The **Rarita-Schwinger theory** – well-known theory of spin 3/2. But this theory has a number of **pathologies.**

Generalized Hamiltonian dynamics: **Dirac bracket** instead of **Poisson bracket**

$$
[F(\vec{x}), G(\vec{y})]_D = [F(\vec{x}), G(\vec{y})] - \int d^3w d^3z [F(\vec{x}), \chi^{\dagger}(\vec{w})] M^{-1}(\vec{w}, \vec{z}) [\chi(\vec{z}), G(\vec{y})]
$$
  

$$
M(\vec{x}, \vec{y}) = [\chi(\vec{x}), \chi^{\dagger}(\vec{y})]
$$

There is **singularity** in a Dirac bracket in weak gauge field limit for RS-theory! **Doesn't allow to construct perturbation theory!**

Solved in **[Stephen L. Adler. Phys. Rev. D, 97(4):045014, 2018]** by introducing of interaction with additional spin ½ field:

$$
S = \int d^4x \left( -\varepsilon^{\lambda\rho\mu\nu} \bar{\psi}_{\lambda} \gamma_5 \gamma_\mu \partial_\nu \psi_\rho + i \bar{\lambda} \gamma^\mu \partial_\mu \lambda - im \bar{\lambda} \gamma^\mu \psi_\mu + im \bar{\psi}_\mu \gamma^\mu \lambda \right)
$$

### Rarita-Schwinger-Adler model of spin 3/2

**•** The interaction **shifts the pole** in the Dirac bracket!

$$
[\Psi_i(\vec{x}), \Psi_j^{\dagger}(\vec{y})]_D = -i \left[ (\delta_{ij} - \frac{1}{2}\sigma_i \sigma_j) \delta^3(\vec{x} - \vec{y}) - \overrightarrow{D}_{\vec{x}} i \frac{\delta^3(\vec{x} - \vec{y})}{m^2 + g \vec{\sigma} \cdot \vec{B}(\vec{x})} \overleftarrow{D}_{\vec{y}} j \right]
$$

Contribution of interaction with an additional field

The **stress-energy tensor** can be obtained by varying with respect to the metric

$$
T^{\mu\nu} = -\frac{2}{\sqrt{-g}} \frac{\delta S}{\delta g_{\mu\nu}}
$$

$$
T^{\mu\nu} = \frac{1}{2} \varepsilon^{\lambda\nu\beta\rho} \bar{\psi}_{\lambda} \gamma_5 \gamma^{\mu} \partial_{\beta} \psi_{\rho} + \frac{1}{8} \partial_{\eta} \Big( \varepsilon^{\lambda\alpha\nu\rho} \bar{\psi}_{\lambda} \gamma_5 \gamma_{\alpha} [\gamma^{\eta}, \gamma^{\mu}] \psi_{\rho} \Big) + \frac{i}{4} \Big( \bar{\lambda} \gamma^{\nu} \partial^{\mu} \lambda - \partial^{\mu} \bar{\lambda} \gamma^{\nu} \lambda \Big) + \frac{i}{2} m \Big( \bar{\psi}^{\mu} \gamma^{\nu} \lambda - \bar{\lambda} \gamma^{\nu} \psi^{\mu} \Big) + (\mu \leftrightarrow \nu). \qquad \text{Traceless unlike the usual} \qquad T^{\mu}_{\mu} = 0 \text{Rarita-Schwinger field}
$$

● The currents can be obtained from Noether's theorem. The **axial current** can be constructed for the  $U(1)_A$  transformation:

$$
j^\mu_A = -i \varepsilon^{\lambda \rho \nu \mu} \bar{\psi}_\lambda \gamma_\nu \psi_\rho + \bar{\lambda} \gamma_\mu \gamma_5 \lambda
$$

#### Chiral anomaly in RSA theory: gauge part

- Since the problem with the Dirac bracket is solved **perturbation theory** can be constructed
- The **chiral (gauge) quantum anomaly** was obtained by the shift method:

**[Stephen L. Adler. Phys. Rev. D, 97(4):045014, 2018]**

**[S. L. Adler, P. Pais, Phys. Rev. D 99, 095037 (2019)]** see also

$$
\langle \partial_{\mu} \hat{j}_{A}^{\mu} \rangle = -\frac{5}{16\pi^{2}} \varepsilon^{\mu\nu\alpha\beta} F_{\mu\nu} F_{\alpha\beta}
$$

Also by the method of conformal three-point functions:

**[G. Y. Prokhorov, O. V. Teryaev, V. I. Zakharov, Phys. Rev. D 106 (2) (2022) 025022]**

● The factor "**5**" differs from what is expected according to the prediction "**3**" based on supergravity **[M. J. Duff, in First School on Supergravity (1982) arXiv:1201.0386]**

However the correspondence is restored if we take into account that there are two additional degrees of freedom with spin  $\frac{1}{2}$ : then *5=3+2*

#### Gravitational chiral anomaly in RSA theory

For a conformally symmetric theory, if

$$
\left\{\n\begin{aligned}\n\frac{\partial_{\mu} T^{\mu\nu}}{\partial_{\mu} T^{\mu\nu}} &= 0, \quad \partial_{\mu} j^{\mu}_{V} = 0, \quad \partial_{\mu} j^{\mu}_{A} = 0,\n\end{aligned}\n\right.
$$
\n
$$
\left\{\n\begin{aligned}\nT^{\mu}_{\mu} &= 0, \quad T_{\mu\nu} = T_{\nu\mu}.\n\end{aligned}\n\right.
$$

It is proven in [J. Erdmenger.Nucl. Phys. B, 562:315–329, 1999], that the three-point function  $\langle T\hat{T}_{\mu\nu}(x)\hat{T}_{\sigma\rho}(y)\hat{j}_{\omega}^{A}(z)\rangle_{c}$  has the **universal form**:

$$
\langle T\hat{T}^{\mu\nu}(x)\hat{T}^{\sigma\rho}(y)\hat{j}^{\omega}_A(z)\rangle_c = \frac{1}{(x-z)^8(y-z)^8}
$$

$$
\times \mathscr{I}^{\mu\nu,\mu'\nu'}_T(x-z)\mathscr{I}^{\sigma\rho,\sigma'\rho'}_T(y-z)t^{TTA}_{\mu'\nu'\sigma'\rho'}(Z)
$$

 $\mathscr{I}^{T}_{\mu\nu,\sigma\rho}(x) = \mathscr{E}^{T}_{\mu\nu,\alpha\beta} I^{\alpha}_{\sigma}(x) I^{\beta}_{\rho}(x) ,$ 

where the notations are introduced:

Introduced:

\n
$$
\mathscr{E}_{\mu\nu,\alpha\beta}^{T} = \frac{1}{2} (\eta_{\mu\alpha}\eta_{\nu\beta} + \eta_{\mu\beta}\eta_{\nu\alpha}) - \frac{1}{4} \eta_{\mu\nu}\eta_{\alpha\beta},
$$
\n
$$
\mathscr{E}_{\mu\nu\sigma\rho\omega}^{T}(Z) = \frac{\mathscr{A}}{Z^{6}} (\mathscr{E}_{\mu\nu,\eta}^{T} \mathscr{E}_{\sigma\rho,\kappa\varepsilon}^{T} \mathscr{E}_{\omega}^{\eta\kappa\lambda} Z_{\lambda})
$$
\n
$$
-6 \mathscr{E}_{\mu\nu,\eta\gamma}^{T} \mathscr{E}_{\sigma\rho,\kappa\delta}^{T} \mathscr{E}_{\omega}^{\eta\kappa\lambda} Z^{\gamma} Z^{\delta} Z_{\lambda} Z^{-1}
$$

#### Gravitational chiral anomaly in RSA theory

Summing 9 correlates (contributions of different), we will obtain: **Matches the form we want!**

$$
\langle T\hat{T}^{\mu\nu}(x)\hat{T}^{\sigma\rho}(y)\hat{j}_{A}^{\omega}(z)\rangle_{c} = -19\left(4\pi^{6} |(x-y)^{5}\right)
$$
  
× $(x-z)^{3}(y-z)^{3}\bigg)^{-1}e_{\vartheta}\bigg(\eta^{\nu\rho}\varepsilon^{\sigma\vartheta\mu\omega} + \eta^{\mu\rho}\varepsilon^{\sigma\vartheta\nu\omega}\bigg)$   
+ $\eta^{\nu\sigma}\varepsilon^{\vartheta\mu\rho\omega} + \eta^{\mu\sigma}\varepsilon^{\vartheta\nu\rho\omega} - 6e^{2}(e^{\nu}e^{\rho}\varepsilon^{\sigma\vartheta\mu\omega}\bigg)$   
+ $e^{\mu}e^{\rho}\varepsilon^{\sigma\vartheta\nu\omega} + e^{\sigma}e^{\nu}\varepsilon^{\vartheta\mu\rho\omega} + e^{\sigma}e^{\mu}\varepsilon^{\vartheta\nu\rho\omega})\bigg)$ 

**[G. Y. Prokhorov, O. V. Teryaev, V. I. Zakharov, Phys. Rev. D 106 (2) (2022) 025022]**

$$
\langle T\hat{T}^{\mu\nu}(x)\hat{T}^{\sigma\rho}(y)\hat{j}^{\omega}_{A}(z)\rangle = \mathscr{A}\left(4(x-y)^{5}\n\times(x-z)^{3}(y-z)^{3}\right)^{-1}e_{\vartheta}\left(\eta^{\nu\rho}\varepsilon^{\sigma\vartheta\mu\omega} + \eta^{\mu\rho}\varepsilon^{\sigma\vartheta\nu\omega}\n+ \eta^{\nu\sigma}\varepsilon^{\vartheta\mu\rho\omega} + \eta^{\mu\sigma}\varepsilon^{\vartheta\nu\rho\omega} - 6e^{2}(e^{\nu}e^{\rho}\varepsilon^{\sigma\vartheta\mu\omega}\n+ e^{\mu}e^{\rho}\varepsilon^{\sigma\vartheta\nu\omega} + e^{\sigma}e^{\nu}\varepsilon^{\vartheta\mu\rho\omega} + e^{\sigma}e^{\mu}\varepsilon^{\vartheta\nu\rho\omega})\right)
$$

(points on the same axis)

We can determine the factor in the anomaly:

$$
\mathscr{A}_{RSA}=-19\mathscr{A}_{s=1/2}=-\frac{19}{\pi^6}
$$

$$
\langle \nabla_{\mu} \hat{j}_{A}^{\mu} \rangle_{RSA} = \frac{-19}{384\pi^2 \sqrt{-g}} \varepsilon^{\mu\nu\rho\sigma} R_{\mu\nu\kappa\lambda} R_{\rho\sigma}{}^{\kappa\lambda}
$$

**–19 times** different from the anomaly for spin  $\frac{1}{2}$ 

#### Gravitational chiral anomaly in RSA theory

#### ● **How to explain the factor -19?**

• How does it **relate** to **previous** calculations?

$$
\langle \nabla_{\mu} \hat{j}_{A}^{\mu} \rangle_{RS} = \frac{-21}{384\pi^2 \sqrt{-g}} \varepsilon^{\mu\nu\rho\sigma} R_{\mu\nu\kappa\lambda} R_{\rho\sigma}^{\quad \kappa\lambda}
$$
  
[M. J. Duff, 1201.0386] [J.J.M. Carrasco, et al. JHEP, 07:029, 2013]

"ghostless" contribution [J.J.M. Carrasco, et al. JHEP, 07:029, 2013]  $-19 = -20 +$ Contribution of spin 1/2  $-19 = -21 + 2$ 

#### Experiment: few words

In non-central collisions of heavy ions, huge magnetic fields and huge angular momentum arise. Differential rotation - different at different points: **vorticity** and **vortices**.



#### Experiment: few words

Vorticity transforms into polarization



**[Nature 548 (2017) 62-65 arXiv:1701.06657 [nucl-ex]]**

- Generation of **hyperon polarization**.
- Both **vorticity** and **acceleration** are essential for polarization.
- Also described based on **Chiral Vortical Effect** (**CVE**) **[Rogachevsky, Sorin, Teryaev, Phys.Rev.C82(2010) 054910],**

**[Baznat, Gudima, Sorin, Teryaev, Phys.Rev.C93, no.3,031902 (2016)]**

$$
\text{CVE:} \quad \langle j_\mu^5 \rangle = \Big(\frac{T^2}{6} + \frac{\mu^2}{2\pi^2}\Big)\omega_\mu
$$

– **Qualitative and quantitative correspondence!**

– Polarization from quantum anomaly  $\sim$  spin crisis and gluon anomaly: **[Efremov, Soffer, Teryaev,** 

**Nucl.Phys.B 346 (1990) 97-114]**

proton spin  $\rightarrow$  hyperon polarization. gluon field  $\rightarrow$  chemical potential\*4-velocity

Also described on the basis of a thermodynamic approach (Wigner function):

**[I. Karpenko, F. Becattini, Nucl.Phys.A 982 (2019) 519- 522]**

#### Experiment: few words

- Is it possible to observe KVE in experiment?
- Is it possible to observe a gravitational chiral anomaly in the hydrodynamics of the matter, produced in heavy ion collisions?
- To observe cubic terms, it is necessary that the gradients (acceleration, vorticity, magnetic field...) give a contribution at least of the order of temperature

 $\omega, a \sim (0.1 - 2)T$ 

**[A. Zinchenko, A. Sorin, O. Teryaev, M. Baznat, J.Phys.Conf.Ser. 1435 (2020)] [F. Becattini, M. Buzzegoli, G. Inghirami, I. Karpenko, A. Palermo, Phys.Rev.Lett. 127 (2021) 27, 272302]**

However, the cubic terms are suppressed by the numerical factor

**KVE:** 
$$
j_{A,S=1/2}^{\mu} = \left(-\frac{1}{24\pi^2}\omega^2 - \frac{1}{8\pi^2}a^2\right)\omega^{\mu}
$$

The **good** news: for spin 3/2 it is enhanced by cubic growth with spin:

The **bad** news: should be suppressed by mass  $e^{-m/T}$  (omega baryon is heavy).

**Idea:** consider massless **quasiparticles** with spin 3/2 in semymetals?

**[I. Boettcher, Phys. Rev. Lett. 124, 127602 (2020)]**

#### Modern development and the problem

A whole set of similar effects has been found at the intersection of quantum field theory and hydrodynamics **[M. N. Chernodub et al. 2110.05471]**.

Experimental search is in progress:

- CME not yet found in ion collisions at RHIC **[D.E. Kharzeev et al. 2205.00120]**.
- Condensed matter copies of the effects are found in semimetals

**[Qiang Li et al. Nature Phys. 12 (2016)]**.