Early universe holographic first order phase transition within Composite Higgs boson model

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AdS/CFT correspondence

Perturbation theory

 $\langle \phi \dots \phi \rangle = \langle \phi_0 \dots \phi_0 \rangle + \lambda \langle \phi_1 \dots \phi_1 \rangle + \lambda^2 \langle \phi_2 \dots \phi_2 \rangle + \dots +$ non-analytical?+solitons? QFT ^{renormalization} CFT

AdS/CFT correspondence

Correlators within AdS/CFT

Solutions of EoM:
$$
\frac{\delta S_{AdS}[\psi]}{\delta \psi} = 0
$$

Near the conformal boundary: $\psi(x, z) \frac{\partial A dS}{z \to 0} z^{d-\Delta} \psi_0(x) + z^{\Delta} \psi_1(x)$
CFT sources: $J \sim \psi_0(x)$ with weight $d - \Delta$

Solutions:
$$
\Rightarrow \mathcal{Z}_{AdS}|_{\partial AdS}[\psi_0] = \mathcal{Z}_{CFT}[J] = \int \mathcal{D}[^{\star} \text{fields of CFT}^{\star}] e^{-S_{CFT}-\mathcal{O} \cdot J}
$$

$$
G_n = \langle \mathcal{O} \dots \mathcal{O} \rangle = \left(\frac{\delta}{\delta J}\right)^n \log \mathcal{Z}[J]\Big|_{J=0} = -\left(\frac{\delta}{\delta \psi_0}\right)^n S_{AdS}[\psi] = -\left(\frac{\delta^n S_{AdS}^{\text{bulk}}}{\delta \psi_0^n}\right) - \left(\frac{\delta^n S_{\partial AdS}^{\text{border}}}{\delta \psi_0^n}\right)
$$

$$
= 0 \text{ due to EOM}
$$

[Some common things about AdS/CFT](#page-1-0) 2/8

Electroweak baryogenesis (Motivation)

Baryon asymmetry problem – matter more than anti-matter

Sakharov's conditions \Leftrightarrow first order phase transition (i.e. CPT-violation)

 $SU(2)_L \times U(1)_Y \rightarrow U(1)_{em}$ is a crossover (not a PT); BSM physics?

[Motivation](#page-3-0) 3/12

Composite Higgs model

 $\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{CH}} + \mathcal{L}_{\text{Int.}}, \quad \mathcal{L}_{\text{CH}}$ – strongly coupled with \mathcal{G} inner symmetry $\left(G \text{ invariant} \atop \text{vacuum}\right) \xrightarrow{\text{spontaneous}}$ breaking $\left(\mathcal{H} \right.$ invariane $\left(\mathcal{H} \right) \Rightarrow$ Goldstone bosons \Rightarrow Higgs boson
phase transition phase transition

$$
\Sigma_{IJ} = \langle \bar{\Psi}_I \Psi_J \rangle = \xi^\top \left[\begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \varsigma \end{pmatrix} + \eta_i \tilde{T}_i \right] \xi \xrightarrow[\text{low energy}]{} \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \varsigma \end{pmatrix} \Rightarrow \begin{array}{c} \text{symmetry} \\ \text{breaking} \end{array}
$$

 Σ_{IJ} is a condensate of the SO(5)-inn.sym. fundamental fields Ψ ; ξ is NB-bosons, η is "radial" fluctuations, ζ is background field

$$
\text{Holographic Model} \atop \text{Dual theory:} \quad S_{AdS}=S_{grav+\phi}+S_X+S_{gauge}+S_{``SM''}+S_{int. \text{``SM''+X}} \xrightarrow{AdS/CFT} S_{SM+CH}
$$

 $S_{\text{grav.}+\phi}$, S_{X} and S_{gauge} can have a PT $\overline{}$ $\overline{}$ \overline{a} $S_{\text{``SM''}}$ and $S_{\text{int. ''SM'' + X}}$ shouldn't have a PT

Matter sector:
$$
S_{X} = \frac{1}{k_{s}} \int d^{5}x \sqrt{|g|} e^{\phi} \left[\frac{1}{2} g^{ab} \operatorname{Tr} \left(\nabla_{a} X^{T} \nabla_{b} X \right) - V_{X}(X) \right]
$$

\n
$$
V_{X}(X) = \operatorname{Tr} \left(-\frac{3}{2L^{2}} X^{T} X - \alpha (X^{T} X)^{2} + \beta (X^{T} X)^{3} + \mathcal{O}(X^{8}) \right)
$$

\n
$$
X_{IJ} \propto \frac{\sqrt{N}}{2\pi} J_{IJ} z + \frac{2\pi}{\sqrt{N}} \Sigma_{IJ} z^{3} + \dots, \quad J_{IJ} \text{ are sources for CH condensate } \Sigma_{IJ}
$$

\n**Our contribution:** (and the following for the matter sector)
$$
V_{\text{eff. CH}}[\Sigma_{IJ}]|_{\text{extrema}} \xrightarrow{\text{AdS/CFT}} S_{X}[X_{IJ}; g_{ab}^{\text{fixed}}, \phi^{\text{fixed}}] |_{J_{IJ}=0}^{\partial \text{AdS}}
$$

Our contribution:	$V_{\text{eff. CH}}[\Sigma_{IJ}] _{\text{extrema}}$	$\frac{\text{AdS/CF1}}{\text{AdS/CF1}}$	$S_X[X_{IJ}; g_{ab}^{\text{fixed}}, \phi^{\text{fixed}}] _{J_{IJ}=0}^{\partial \text{AdS}}$		
Section of matter	4-dimensional	4-dimensional	4-dimensional	4-dimensional	4-dimensional

Phase transition

$$
\frac{\partial V_{\text{eff}}^{\text{CH}}}{\partial \langle \varphi \rangle}\Big|_{\langle \varphi \rangle_0}=0 \quad \Rightarrow \quad \Big(\langle \varphi \rangle_0,V_{\text{eff}}\big|_{\langle \varphi \rangle_0}\Big)
$$

Effective potential extremal values and the positions allow one to judge about PT:

- **t** trivial minimum (vacuum) only ⇒ there is no PT;
- **non-trivial true vacuum with the potential barrier** ⇒ **1-st PT**;
- \blacksquare non-trivial true vacuum without a potential barrier \Rightarrow there is no PT.

The extrema of the effective quantum potential V_{eff} : T is the plasma temperature, $\langle \varphi \rangle$ is the vacuum expectation.

[PT and GW](#page-6-0) 6/21 Unscaled schematic illustration! Data in real scale are Phys. Rev. D 108, 115011.

Observations

Picks of the GW spectrum estimated within Holographic Composed Higgs model. There is only the scalar part produced during initial collisions of the bubble walls (i.e. sound and turbulence contributions are not currently included in the rough estimate.)

[PT and GW](#page-6-0) 7/26

Large D limit

$$
\text{Dual theory:}\quad \mathsf{S}_{\text{AdS}}=\mathsf{S}_{\text{grav.}^+\phi}+\mathsf{S}_X+\mathsf{S}_{\text{gauge}}+\mathsf{S}_{\text{``SM''}}+\mathsf{S}_{\text{int.}}\text{``SM''}\text{+}\mathsf{x}
$$

 S_X is considered (without fluctuations); $S_{grav,t\phi}$ and S_{gauge} are left

$$
S_{\text{grav.}+\phi} = \frac{1}{l_p^3} \int d^5 x \sqrt{|g|} e^{2\phi} \Big[-R + 2|\Lambda| - 4g^{ab} \partial_a \phi \partial_b \phi - V_{\phi}(\phi) \Big]
$$

$$
S_{\text{grav.}+\phi} \text{ within large D limit: } g_{ab} = g_{ab}^{(0)} + \frac{1}{D}g_{ab}^{(1)} + \frac{1}{D^2}g_{ab}^{(2)} + \dots \text{ [Emparan'20]}
$$

near-AdS_d (Poincaré patch):
$$
A(z) = 1 + o(e^D)
$$
, $B(z) = 1 + o(e^D)$, $f(z) = 1 - \left(\frac{z}{z_H}\right)^{D-1} + \mathcal{O}(e^D)$

$$
ds^2 = A(z)\left(-f(z)dt^2 + \frac{dz^2}{f(z)} + B(z) d\Omega_{D-2}\right) \xrightarrow{D \to \infty} \frac{l^2}{D^2} \left(-y^2 d\tau^2 + dy^2\right) + o(e^{-D}) d\Omega_{D-2}
$$

[Sector of gravity](#page-8-0) 8/37 The simplest case for dilaton in AdS $\quad \phi'' - (\phi')^2 + m^2(\phi^2 + 2\phi) = 0 \quad$ no Lie symmetries

Large D limit for black branes and GHY

For the black brane with normal n^A , tangent velocity u^A and extrinsic curvature \mathcal{K}_{AB}

$$
\left(\frac{\nabla^2 u_A}{K} - \frac{\nabla_A K}{K} + u^A K_{AB} - u^B \nabla_B u^A\right) \mathcal{P}_C^A = \mathcal{O}\left(\frac{1}{D}\right), \quad \mathcal{P}^{AB} = \delta^{AB} + u^A u^B
$$

for "flat" metric η_{AB} (i.e. without any horizon) [Bhattacharyya'16]

$$
g_{AB} = \eta_{AB} + \frac{(n^A - u^A)(n^B - u^B)}{\psi^{D-3}} + \mathcal{O}\left(\frac{1}{D}\right), \quad \psi\big|_{\text{horizon}} = 1 \text{ is a scalar function}
$$

$$
S_{grav.} = \underbrace{\int_{\mathcal{M}} R dVol_D}_{EH with out BH} + \underbrace{\int_{\partial \mathcal{M}} K dVol_{D-1}}_{with PT}
$$
\n
$$
\underbrace{\frac{1}{2} S_{matter}}_{with PT}
$$
\n
$$
\underbrace{\frac{1}{2} S_{matter}}_{S/2} = \frac{1}{2} S_{matter} \underbrace{\frac{1}{2} S_{\text{inter}} \underbrace{\frac{1}{2} S_{\text{inter}}}{\frac{1}{2} S_{\text{inter}}}
$$

[Sector of gravity](#page-8-0) and the sector of gravity $9/37$

Effective CFT within large D limit

$$
\underbrace{\int \frac{d^{d+1}x\sqrt{-g}}{16\pi G_{d+1}} \left(R + \frac{d(d-1)}{L^2} - \frac{L^2}{g_s^2} F^2 \right)}_{\text{Einstein-Maxwell RNAdS}_{d+1}} + \underbrace{\int \frac{d^2x\sqrt{-he^{-2\psi}}}{16\pi G_2} \left(R_2 - \frac{1}{g_s^2} F^2 + 4(\nabla\psi)^2 + \frac{1}{L^2} \right)}_{\text{Einstein-Maxwell-dilaton RNAdS}_2}
$$
\n
$$
T^{ab} \Big|_{\partial RN \text{-AdS}_2} = \frac{2}{\sqrt{-h}} \underbrace{\frac{\delta S}{\delta h_{ab}}}_{\delta h_{ab}}; \text{ asymptotic diffeomorphism}_{\text{(gauge invariant)}} \chi^\mu \to \chi^\mu + \epsilon^\mu
$$
\n
$$
(\delta_\epsilon + \delta_\Lambda) T_{tt} = \underset{\text{transformation}}{\text{unsformation + }\atop \text{term}} + \underset{\text{term}}{\text{anomalous}}, \quad A_\mu \to A_\mu + \partial_\mu \Lambda
$$
\n
$$
\text{anomalous} \implies c = \frac{3}{2\pi G_2} = \frac{3\Omega_{d-1}}{2\pi G_{d+1}} r_0^{d-1} \implies s_{\text{CFT}} = \frac{\pi^3}{3} T c = \frac{1}{4G_2} \quad \begin{pmatrix} \text{by Cardy} \\ \text{formula} \end{pmatrix}
$$
\n
$$
\text{Bekenstein-Hawking for RN-AdS}_{d+1} \quad s = \frac{\pi^3}{3} T c = \frac{1}{4G_2} = \frac{\text{Area}}{G_{d+1}} \quad \text{[Guo'16]}
$$
\n
$$
\text{Sector of gravity}
$$

Hawking-Page PT within large D limit eff. CFT

something that looks like a conclusion

 $F_{\text{thermal AdS}} = -pV \xrightarrow{\text{Hawking-Page}} F_{\text{BH}} = M - TS - \mu Q - pV$

- 1. $S_{\phi + \text{grav}} + S_{\text{gauge}}$ asymptotic RN-AdS_{d+1} (unknown) solutions (or "near-AdS×any-compact" for top-down approach),
- 2. $\xrightarrow{\text{large D}}$ RN-AdS₂ (RN-AdS₃) solutions in leading 1/*d* order,

3.
$$
\xrightarrow[\text{asymptotic}]{\text{border}} \rightarrow \text{central charge of CFT}_1 \text{ (CFT}_2),
$$

- 4. $\frac{Cardy's}{formals}$ free energy and Hawking-Page phase transition (in RN-AdS_{d+1}), formula
- 4. $\frac{1}{\text{formula}}$ free energy and Hawking-Page phase transition (in
5. $\frac{\text{AdS/CFT}}{\text{600}}$ confinement/deconfinement-like PT for CH model (as it works for AdS/QCD).

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Thank you for your attention!

Phys. Rev. D 108, 115011 (matter sector only)

System of units:
$$
1 = c = \hbar = k_B = \pi = e = i = -1 = 2
$$
kg

Holographic model

 \mathcal{L}_{CH} – strongly coupled \Rightarrow consider $N \gg 1 \Rightarrow \mathcal{Z}_{CH}[J] = \mathcal{Z}_{AdS}[J]$ The dual theory: $\mathcal{Z}_{\sf AdS}[J]\sim \exp\Big(-S_{\sf AdS}[J]\Big)$ is weakly coupled $\;\Rightarrow\;$ quasiclassical limit

The asymptotic behavior near the conformal border ∂AdS of the dual theory fields **defines the sources** of the CH operators (i.e. the **correlator functions**)

$$
X_{IJ} \stackrel{z \to 0}{\longrightarrow} \frac{\sqrt{N}}{2\pi} J_{IJ} z + \frac{2\pi}{\sqrt{N}} \Sigma_{IJ} z^3 + \dots \quad X_{IJ} : \text{AdS}_5 \stackrel{\text{dual}}{\iff} \Sigma_{IJ} : \mathbb{R}^{1,3}
$$

Holography is the **duality** between **strongly coupled** theory on the border and **weakly coupled** (quasiclassical) bulk theory.

 $F = -T \log \mathcal{Z}_{CH} \sim TS_{AdS} \propto Vol_4 \cdot \mathcal{F}$ In homogeneous case $(\chi = \chi(z)) : \mathcal{F} \propto V_{eff}[\chi]$

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Finite temperature

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Action of the holographic model

$$
S_{\text{tot}} = S_{\text{grav}+\phi} + S_{\text{X}} + S_{\text{A}} + S_{\text{SM}} + S_{\text{int}}, \quad S_{\text{A}} = -\frac{1}{g_5^2} \int d^5 x \sqrt{|g|} e^{\phi} g^{\alpha c} g^{\beta d} F_{\alpha b} F_{\text{cd}}
$$

\n
$$
S_{\text{grav}+\phi} = \frac{1}{l_p^3} \int d^5 x \sqrt{|g|} e^{2\phi} \Big[-R + 2|\Lambda| - 4g^{\alpha b} \partial_{\alpha} \phi \partial_{b} \phi - V_{\phi}(\phi) \Big], \quad a, b = 0, \dots 4
$$

\n
$$
S_{\text{int}} = \epsilon^4 \int_{z=\epsilon} d^4 x \sqrt{|g^{(4)}|} \Big[c_Y B_{\mu} \operatorname{Tr} (T_Y A^{\mu}) + c_W W_{k,\mu} \operatorname{Tr} (T_k A^{\mu}) + \mathcal{L}_{\psi} \Big]
$$

\n
$$
S_{\text{X}} = \frac{1}{k_s} \int d^5 x \sqrt{|g|} e^{\phi} \Big[\frac{1}{2} g^{\alpha b} \operatorname{Tr} (\nabla_{\alpha} X^T \nabla_{b} X) - V_X(X) \Big], \quad \nabla_{\alpha} X = \partial_{\alpha} X + [A_{\alpha}, X], \quad A_{\alpha} = 0
$$

\n
$$
V_X(X) = \operatorname{Tr} \Big(-\frac{3}{2L^2} X^T X - \frac{\alpha}{4} (X^T X)^2 + L^2 \frac{\beta}{6} (X^T X)^3 + O(X^8) \Big)
$$

\n
$$
L \cdot X_{IJ} \sim \frac{\sqrt{N}}{2\pi} J_{IJ} \tilde{z} + \frac{2\pi}{\sqrt{N}} \Sigma_{IJ} \tilde{z}^3 + \dots
$$

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Geometry

$$
S_{\text{grav}+\phi} = \frac{1}{l_p^3} \int d^5 x \sqrt{|g|} e^{2\phi} \Big[-R + 2|\Lambda| - 4g^{ab} \partial_a \phi \partial_b \phi - V_{\phi}(\phi) \Big], \quad a, b = 0, \dots 4
$$

$$
ds^{2} = \frac{L^{2}}{\tilde{z}^{2}} A(\tilde{z})^{2} \left(f(\tilde{z}) d\tau^{2} + \frac{d\tilde{z}^{2}}{f(\tilde{z})} + d\vec{x}^{2} \right), \quad \phi = \phi(\tilde{z})
$$

$$
f = 1 - \frac{\tilde{z}^{4}}{z_{H}^{4}}, \quad \phi = \tilde{\phi}_{2} \tilde{z}^{2}, \quad z_{H} = \frac{1}{\pi T}.
$$

Effective field theory

$$
\mathcal{Z}[J] = \int \mathcal{D}\phi \, e^{-S-J\cdot\phi} =: e^{W[J]}
$$

$$
\Gamma[\langle\phi\rangle] = W[J] - \frac{\delta W[J]}{\delta J} \cdot J = \int_X d^d x \Big(\underbrace{K_{\text{eff}}[\partial\langle\phi\rangle]}_{=0 \text{ if } \langle\phi\rangle = \text{const}} + V_{\text{eff}}[\langle\phi\rangle] \Big) \quad - \text{effective action}
$$

$$
\text{Effective potential:} \quad V_{\text{eff}} = \frac{1}{\text{Vol}_4} \Gamma
$$

Equation of motion (EoM):

$$
\begin{aligned}\n\text{fective potential:} \quad V_{\text{eff}} &= \frac{1}{\text{Vol}_4} \Gamma \\
\frac{\delta \Gamma}{\delta \langle \phi \rangle} &= J \xrightarrow{\langle \phi \rangle = \text{const}} \frac{\delta V_{\text{eff}}}{\delta \langle \phi \rangle} \xrightarrow{J=0} 0 \quad \text{gives extrema condition}\n\end{aligned}
$$

[Effective theory approach](#page-17-0) in the state of the state o

Effective potential

Extrema condition:
$$
V_{\text{eff}}|_{\text{extrema}} = V_{\text{eff}}|_{J=0} \Leftrightarrow G_0;
$$

\nAdS/CFT: $G_0 \Leftarrow$ boundary term of dual theory $S_{\partial AdS}$

\nVolx $V_{\text{eff}}|_{\text{extrema}} = G_0 = W[J = 0]$

\n $\frac{\text{AdS/CFT}}{\text{AdS}} \int_{\partial AdS}^{\psi_0 = (0, \text{det}(S_0))} S_{\text{AdS}} \left| \frac{\psi_0 = (0, \text{det}(S_0))}{\partial A} \right|_{\partial AdS}$

$$
Vol_X V_{eff}|_{extrema} = G_0 = W[J = 0] \stackrel{\text{AdS/CFT}}{\longrightarrow} S_{\text{AdS}}|_{\partial \text{AdS}}^{\psi_0 = 0}
$$

Extrema condition & duality: $J = \psi_0 = 0$; duality: $\langle \phi \rangle = \psi_1$ δV_{eff} $\delta\langle\phi\rangle$ $\frac{\text{Add}(CFT)}{\text{Ind}(CFT)}$ AdS/CFT δ $\delta \psi_1$ $(S[\psi]|_{\partial\mathsf{AdS}})\Big|_{\psi_0=0}=0$ (with assumption) gives vacuum expectation values: $\;\;\{\langle \phi \rangle_{\sf min\,1},\langle \phi \rangle_{\sf min\,2},\dots\}\;\;\;$ $-$ possible vacuums

Extrema positions and values $\left\{ \left(\langle\phi\rangle_{\mathsf{min}}}, V_{\mathsf{eff}}[\langle\phi\rangle_{\mathsf{min}}]\right)\right\} \Rightarrow$ phase transitions

[Effective theory approach](#page-17-0)

Gravitational Waves

The spectrum of the gravitational waves can be estimated as (within the approach of **relativistic** velocity of the bubble walls $v_w \sim 1$)

$$
\Omega_{\rm GW} h^2 = 1.67 \cdot 10^{-5} \kappa \Delta \left(\frac{\beta}{H_*}\right)^{-2} \left(\frac{\alpha}{1+\alpha}\right)^2 \left(\frac{g_*}{100}\right)^{-\frac{1}{3}}
$$

Only scalar waves! Sound waves and turbulence are not included! We **estimate** only scalar waves produced during initial collisions.

$$
f_0 = 1.65 \cdot 10^{-5} \text{Hz} \cdot \frac{f_*}{\beta} \frac{\beta}{H_*} \frac{7}{0.1 \text{TeV}} \left(\frac{g_*}{100}\right)^{\frac{1}{6}} \text{Hz}
$$

 $(\Omega_{\textsf{GW}} h^2, f_0)$ -curve is the estimation GW amplitude (peak value). It does not contain the spectral shape S (f_{0}) (in this case S $(f_{0}=f_{0}^{\mathsf{peak}})$ j_{0}^{peak}) = 1).

Temperature estimations

Experimental restrictions ⇔ mass of the **lightest predicted** particle.

rimental restrictions
$$
\Leftrightarrow
$$
 mass of the **lightest predicted** pan

$$
\Sigma_{IJ} = \xi^{\top} \left[\begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \zeta \end{pmatrix} + \eta_i \tilde{T}_i \right] \xi \begin{pmatrix} AdS/CFT \\ dual \text{ to } X_J \end{pmatrix} X_{IJ} \rightarrow \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \chi \end{pmatrix}
$$

$$
m_{\eta} \sim m_{\delta \chi} \text{ fluctuation mass} \sim \text{slope of the "hat".}
$$

$$
\chi(z) \to \chi(z) + \delta \chi(t, \vec{x}, z) \quad \Rightarrow \quad \text{EoM}_{z}[\chi] \to \text{EoM}_{t, \vec{x}, z}[\chi + \delta \chi] \quad \frac{\partial \text{AdS}}{\text{diag}} \quad \boxed{2m^2 = \phi_2}
$$

$$
T = \frac{1}{\pi} \frac{1}{z_H} \quad \Rightarrow \quad T = \frac{m}{\pi} \sqrt{\frac{2}{\phi_2}}, \quad z_H^2 = \frac{\phi_2}{2m^2}
$$

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Bubble free energy

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Holographic effective potential

$$
\mathcal{Z}_{CH}[J] = \int \mathcal{D}\varphi \exp\left(-S[\varphi] - \int d^4x \,\varphi(x) J(x)\right) \stackrel{\text{def}}{=} e^{-W[J]}
$$
\n
$$
\langle \varphi \rangle = \frac{\delta W[J]}{\delta J}\Big|_{J=0}, \ \Gamma[\langle \varphi \rangle] = W[J] - \int d^4x \frac{\delta W[J]}{\delta J(x)} J(x) - \text{Effective Action}
$$
\n
$$
\boxed{\text{EoM: } \frac{\delta \Gamma}{\delta \langle \varphi \rangle} = J} \qquad \begin{array}{c} \text{Homogeneous} \\ \text{Solution} \end{array} \Rightarrow \ \langle \varphi \rangle = \text{const}_{\mathbb{R}^{1,3}} \Rightarrow \Gamma = -\text{Vol}_4 V_{\text{eff}} - \begin{array}{c} \text{Effective Action} \\ \text{potential} \end{array}
$$
\n
$$
\mathcal{Z}[J] \xrightarrow{\text{AdS}/\text{CFT}} \mathcal{Z}_{\text{AdS}} \qquad \begin{array}{c} \text{quasiclassical} \\ \text{approximation} \end{array} e^{-S_{\text{AdS}}}\Big|_{\partial \text{AdS}} - \text{quasiclassical} \\ \text{non-perturbative} \\ \boxed{V_{\text{eff}} = -\frac{1}{\text{Vol}_4} S_{\text{AdS}}}\Big|_{\partial \text{AdS}} - \text{boundary term of the bulk theory} \\ \text{defines quantum effective potential} \end{array}
$$

[Effective theory approach](#page-17-0)

$$
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$$

"Extrema" curves

$$
\frac{\delta S_{\chi}}{\delta \chi} = 0 \Rightarrow \chi \xrightarrow{z \to 0} Jz + \left(\sigma - \left(\frac{3}{2}J^3 + \phi_2 J\right) \log z\right) z^3 + o(z^5) - \text{give the sources}
$$

Knowing the extrema of the effective potential and its values at these points, we can judge abut the phase transition

$$
V_{\text{eff}} = -\frac{1}{\text{Vol}_{4}} S_{\chi} \Big|_{\partial \text{AdS}} \Rightarrow \text{ for effective : } \text{Vol}_{4} \frac{\delta V_{\text{eff}}}{\delta \langle \varphi \rangle} = J \Rightarrow \text{ extrema condition is } \Rightarrow J = 0
$$
\n
$$
\xrightarrow{\text{``extreme''}} \text{solutions}
$$
\n
$$
\xrightarrow{\text{``extreme''}} \text{solutions}
$$
\n
$$
\frac{\text{extrema}}{\delta \langle \varphi \rangle} = 0 \Rightarrow \text{ a new condition for } \phi_{2} \text{ and } \langle \varphi \rangle
$$
\n
$$
T \sim \frac{1}{\sqrt{\phi_{2}}}, \quad \frac{\delta V_{\text{eff}}}{\delta \langle \varphi \rangle} = 0 = \frac{\delta}{\delta \sigma} S_{\chi} \left[\chi_{\text{Sol.}}(z; J, \sigma) \right] \Big|_{J=0} \Rightarrow \quad \{\sigma_{1}, \dots, \sigma_{n}\} \text{ - extrema}
$$

Effective theory apps (soft and to $\langle \varphi \rangle$, vacuum average of the effective theory $140/12$

Nucleation ratio

Baryogenesis generates enough asymmetry (enough efficient) if there is one bubble per Hubble volume

Nucleation:	$AT^4e^{-\frac{F_c}{T}}$	$\sim H^4(T) = \left(\frac{T^2}{M_{Pl}}\right)^4$	Expansion of the Universe
Bubbles produced per time × space volume	$1/(Hubble time × volume)$		

 $F = F[\langle \varphi \rangle, R]$ – Free energy of the bubble; R is the radius of the bubble

Hubble horizon (time, volume, radius) — speed of receding object behind it is greater than the speed of light (Don't confuse with cosmological horizon)

Bubble appears with a certain size. It defines with "micro-physics". If its radius is grater, then critical one $\frac{\partial F}{\partial R}\big|_{R_{\bf C}\triangleq R}$, the bubble grow. Otherwise, it bursts.

It gives $F_{\text{C}} \stackrel{\text{def}}{=} F(R_{\text{C}})$ and defines nucleation ratio and "viability of the model".

[Effective theory approach](#page-17-0)

Estimations of the nucleation ratio

[Effective theory approach](#page-17-0) in the control of the control o

Potentials of CFT and the dual theory

$$
V_{\chi} = a_2 \chi^2 + a_4 \chi^4 + a_6 \chi^6, \quad a_2 < 0, a_4 < 0, a_6 > 0 \quad \text{no barrier}
$$

$$
V_{\text{eff}} = b_2 \langle \varphi \rangle^2 + b_4 \langle \varphi \rangle^4 + b_6 \langle \varphi \rangle^6, \quad b_2 > 0, b_4 < 0, b_6 > 0 \quad \text{there's a barrier}
$$

in details:

- $V_{\text{eff}} = V_{\text{eff}}[\langle \varphi \rangle]$ describes a quantum objects at the border. V_{γ} is a dual classical potential in the bulk.
- $V_{\text{eff}} = -\frac{1}{\text{V}_0}$ $\frac{1}{\text{Vol}_4}$ S_{AdS} includes the solutions of the EoM $\frac{\delta S_\chi}{\delta \chi=0}$ in bulk. In other words, V_{eff} includes physics of AdS

Conditions for the dual theory potential

$$
V_{\chi}(\chi)=\frac{m^2}{2}\chi^2-\frac{D}{4L^2}\lambda\chi^4+\frac{\lambda^2\gamma}{6L^2}\chi^6\text{ is the expansion of a more general theory}
$$

Suggestions:

- The potential V_χ always has true vacuum with E_{\min} $(V_\chi \xrightarrow{\chi\to\pm\infty} \infty).$ So we may use any even power χ^n instead of the last term $\chi^6.$
- **The expansion of** V_v **has certain sign of the second term** $\lambda > 0$ (the first one m^2 chosen for the theory to be conformal in AdS).
- \blacktriangleright Higher orders of the expansion don't give new minima at the considered temperatures.

The certain parametrization has been chosen with respect to the "symmetries"

[Effective theory approach](#page-17-0) 144/12 "Scale invariace", defining the coefficents $L \rightarrow L'$ $\chi \rightarrow$ $\sqrt[3]{\lambda} \chi^{\frac{1}{2}}$ Conformality near the AdS border ("correct" conformal weights) : $\Delta_+ = 1$ ⇒ $m^2 = -\frac{D}{3L}$ $3L^2$ D is for the Large D limit. But its usage doesn't give any results. (to keep interaction constants finite at $D \to \infty$)

SM - CH model interactions

CH model interactions\n
$$
F \frac{\text{thin walls}}{\text{approximation}} 4\pi R^2 \mu - \frac{3\pi}{4} R^3 \left(\mathcal{F}_{\text{out}} - \mathcal{F}_{\text{in}} \right) - \text{physical units are required}
$$
\nFix the Parameters (Interaction with Standard Model - bulk gauge fields)

- Þ. Fix the Parameters (Interaction with Standard Model – bulk gauge fields)
- Physical Units (Infrared Regularization and finite temperature "radial" Þ. heavy fluctuations)

$$
W_{\mu}^{\alpha}J_{L}^{\alpha\,\mu}+B_{\mu}J_{Y}^{\mu}\quad \Leftrightarrow\quad J^{\mu}\sim A^M\,\text{- bulk } \mathcal{G} \text{ gauge field}
$$

The physical values can be estimated without gauge field:

$$
\Sigma_{IJ} = \langle \bar{\Psi}_I \Psi_J \rangle = \xi^\top \left[\begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & X \end{pmatrix} + \eta_i \tilde{T}_i \right] \xi \quad \Leftrightarrow \quad \frac{1}{T} \propto \sqrt{\phi_2} \sim \mu_{IR} \sim m_\eta \gtrsim 10 \text{ TeV}
$$

$$
m_\eta \quad \Leftarrow \quad X \to X + \delta X \text{ - correction of the background field} \quad \Rightarrow \quad \eta \text{ - pNG boson}
$$

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CH gauge field

$$
\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{CH} + B_{\mu} \operatorname{Tr} \left(T_{Y} \hat{J}^{\mu} \right) + W_{k,\mu} \operatorname{Tr} \left(T_{k} \hat{J}^{\mu} \right) + \sum_{r} \bar{\psi}_{r} \mathcal{O}_{r} + \text{h.c.}
$$
\n
$$
= \mathcal{L}_{\text{interactions}}
$$
\n
$$
SO(5) \times U(1) : A_{M} = A_{M}^{K} T^{K} + A_{M,Y} T_{Y}
$$
\n
$$
SO(5) \to SO(4) : \underbrace{A_{M}^{K} T^{K}}_{\in SO(5)} \to \underbrace{A_{M}^{0} T^{0}}_{\in SO(4)} + \underbrace{A_{M}^{i} T^{i}}_{\in SO(5)/SO(4)}
$$
\n
$$
SO(4) \cong SU(2) \times SU(2) : A_{M}^{K} T^{K} = \underbrace{A_{M}^{k, L} T_{L}^{k}}_{\in SU(2)_{L}} + \underbrace{A_{M}^{k, R} T_{R}^{k}}_{\in SU(2)_{R}}
$$
\n
$$
conserved currents: \hat{J}_{\mu} \xleftarrow{dual} A_{\mu}(t, x, z) \Big|_{z=0}^{\partial \text{AdS}}, \text{ holographic gauge: } A_{z} = 0
$$
\n
$$
\mathcal{O} \xleftarrow{dual} \mathcal{J}(A, \Psi, \phi, \dots) - \text{composite operators of the CH fields}
$$

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