# Early universe holographic first order phase transition within Composite Higgs boson model

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# **AdS/CFT correspondence**

#### **Perturbation theory**

$$\langle \phi \dots \phi \rangle = \langle \phi_0 \dots \phi_0 \rangle + \lambda \langle \phi_1 \dots \phi_1 \rangle + \lambda^2 \langle \phi_2 \dots \phi_2 \rangle + \dots + \text{non-analytical?} + \text{solitons?}$$
 QFT 
$$\xrightarrow{\text{renormalization}} \text{CFT}$$

#### AdS/CFT correspondence

strong coupled  $\lambda_{\text{CFT}}\gg 1$  with  $\lambda_{\text{CFT}}\sim \frac{1}{\lambda_{\text{AdS}}}$  dual theory  $\lambda_{\text{AdS}}\ll 1$  weak coupled

# **Correlators within AdS/CFT**

Solutions of EoM: 
$$\frac{\delta S_{AdS}[\psi]}{\delta \psi} = 0$$

Near the conformal boundary: 
$$\psi(x,z) \xrightarrow[z\to 0]{\partial {\sf AdS}} z^{d-\Delta} \psi_0(x) + z^{\Delta} \psi_1(x)$$

CFT sources:  $J \sim \psi_0(x)$  with weight  $d - \Delta$ 

Solutions: 
$$\Rightarrow \mathcal{Z}_{AdS}|_{\partial AdS}[\psi_0] = \mathcal{Z}_{CFT}[J] = \int \mathcal{D}[*fields of CFT*] e^{-S_{CFT}-\mathcal{O} \cdot J}$$

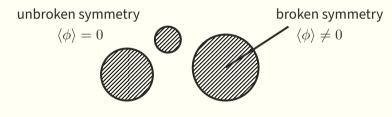
$$G_n = \langle \mathcal{O} \dots \mathcal{O} \rangle = \left(\frac{\delta}{\delta J}\right)^n \log \mathcal{Z}[J]|_{J=0} = -\left(\frac{\delta}{\delta \psi_0}\right)^n S_{AdS}[\psi] = -\left(\frac{\delta^n S_{AdS}^{bulk}}{\delta \psi_0^n}\right) - \left(\frac{\delta^n S_{\partial AdS}^{border}}{\delta \psi_0^n}\right)$$

=0 due to EoM

# **Electroweak baryogenesis (Motivation)**

Baryon asymmetry problem – matter more than anti-matter

Sakharov's conditions  $\Leftrightarrow$  first order phase transition (i.e. CPT-violation)



 ${\rm SU}(2)_{\rm L}\times {\rm U}(1)_{\rm Y} \rightarrow {\rm U}(1)_{\rm em} \quad \text{is a crossover (not a PT); BSM physics?}$ 

Motivation 3/12

# **Composite Higgs model**

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{CH} + \mathcal{L}_{Int.}, \quad \mathcal{L}_{CH} \text{ - strongly coupled with } \mathcal{G} \text{ inner symmetry}$$
 
$$\left( \mathcal{G} \overset{\text{invariant}}{\text{vacuum}} \right) \xrightarrow{\text{breaking}} \left( \mathcal{H} \overset{\text{invariane}}{\text{vacuum}} \right) \Rightarrow \overset{\text{Goldstone bosons } \ni \text{Higgs boson phase transition}$$

$$\Sigma_{\mathit{IJ}} = \langle \bar{\Psi}_{\mathit{I}} \Psi_{\mathit{J}} \rangle = \xi^\top \begin{bmatrix} \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \varsigma \end{pmatrix} + \eta_i \tilde{T}_i \end{bmatrix} \xi \xrightarrow[\text{low energy}]{\text{SO}(5) \to \text{SO}(4)}} \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \varsigma \end{pmatrix} \Rightarrow \text{symmetry breaking}$$

 $\Sigma_{IJ}$  is a condensate of the SO(5)-inn.sym. fundamental fields  $\Psi$ ;  $\xi$  is NB-bosons,  $\eta$  is "radial" fluctuations,  $\varsigma$  is background field

Sector of matter 4/12

#### **Holographic Model**

Dual theory:  $S_{AdS} = S_{grav.+\phi} + S_X + S_{gauge} + S_{"SM"} + S_{int. "SM" + X} \xrightarrow{AdS/CFT} S_{SM+CH}$ 

 $S_{\text{gray},+\phi}$ ,  $S_{\text{X}}$  and  $S_{\text{gauge}}$  can have a PT

S<sub>"SM"</sub> and S<sub>int "SM" + x</sub> shouldn't have a PT

Matter sector: 
$$S_{\rm X} = \frac{1}{k_{\rm S}} \int d^5 x \sqrt{|g|} e^{\phi} \left[ \frac{1}{2} g^{ab} \operatorname{Tr} \left( \nabla_a X^T \nabla_b X \right) - V_X(X) \right]$$
 
$$V_X(X) = \operatorname{Tr} \left( -\frac{3}{2L^2} X^T X - \alpha (X^T X)^2 + \beta (X^T X)^3 + \mathcal{O}(X^8) \right)$$
 
$$X_{IJ} \propto \frac{\sqrt{N}}{2\pi} J_{IJ} z + \frac{2\pi}{\sqrt{N}} \Sigma_{IJ} z^3 + \dots, \quad J_{IJ} \text{ are sources for CH condensate } \Sigma_{IJ}$$

#### Our contribution:

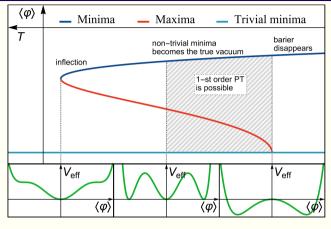
(and the following for the matter sector) 4-dimensional

 $V_{\text{eff. CH}}[\Sigma_{IJ}]\Big|_{\text{extrema}} \stackrel{\text{AdS/CFT}}{=\!=\!=\!=} S_{X}[X_{IJ}; g_{ab}^{\text{fixed}}, \phi^{\text{fixed}}]\Big|_{J_{IJ}=0}^{\partial \text{AdS}}$ 

dual 5-dimensional

 $\approx \frac{\pi^2}{2}/12$ 

#### **Phase transition**



$$\left. \frac{\partial \textit{V}_{\text{eff}}^{\text{CH}}}{\partial \langle \varphi \rangle} \right|_{\langle \varphi \rangle_0} = 0 \quad \Rightarrow \quad \left( \langle \varphi \rangle_0, \textit{V}_{\text{eff}} \right|_{\langle \varphi \rangle_0} \right)$$

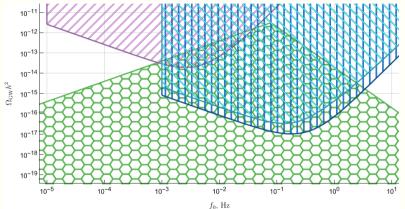
Effective potential extremal values and the positions allow one to judge about PT:

- trivial minimum (vacuum) only ⇒ there is no PT;
- non-trivial true vacuum with the potential barrier  $\Rightarrow$  1-st PT;
- non-trivial true vacuum without a potential barrier ⇒ there is no PT.

The extrema of the effective quantum potential  $V_{\rm eff}$ : T is the plasma temperature,  $\langle \varphi \rangle$  is the vacuum expectation.

Unscaled schematic illustration! Data in real scale are Phys. Rev. D 108, 115011.

#### **Observations**



Picks of the GW spectrum estimated within Holographic Composed Higgs model. There is only the scalar part produced during initial collisions of the bubble walls (i.e. sound and turbulence contributions are not currently included in the rough estimate.)

PT and GW 7/26

#### Large D limit

Dual theory:  $S_{AdS} = S_{grav.+\phi} + S_X + S_{gauge} + S_{"SM"} + S_{int. "SM" + X}$ 

 $S_{\rm X}$  is considered (without fluctuations);  $S_{{\rm grav},+\phi}$  and  $S_{{\rm gauge}}$  are left

$${\cal S}_{
m grav.+\phi} = rac{1}{l_P^3} \int d^5 x \sqrt{|g|} {
m e}^{2\phi} \Big[ - {\it R} + 2 |\Lambda| - 4 g^{ab} \partial_a \phi \partial_b \phi - {\it V}_\phi(\phi) \Big]$$

$$S_{\mathsf{grav},+\phi}$$
 within large D limit:  $g_{ab}=g_{ab}^{(0)}+rac{1}{D}g_{ab}^{(1)}+rac{1}{D^2}g_{ab}^{(2)}+\dots$  [Emparan'20]

near-AdS<sub>d</sub> (Poincaré patch): 
$$A(z) = 1 + o(e^D), B(z) = 1 + o(e^D), f(z) = 1 - \left(\frac{z}{z_H}\right)^{D-1} + \mathcal{O}(e^D)$$

$$ds^2 = A(z) \left( -f(z)dt^2 + \frac{dz^2}{f(z)} + B(z) d\Omega_{D-2} \right) \xrightarrow{D \to \infty} \frac{l^2}{D^2} \left( -y^2 d\tau^2 + dy^2 \right) + o(e^{-D}) d\Omega_{D-2}$$
The simplest case for dilaton in AdS  $\phi'' - (\phi')^2 + m^2(\phi^2 + 2\phi) = 0$  no Lie symmetries

The simplest case for dilator in AdS  $\psi = (\psi) + m (\psi + 2\psi) = 0$  for the symmetries

#### Large D limit for black branes and GHY

For the black brane with normal  $n^A$ , tangent velocity  $u^A$  and extrinsic curvature  $K_{AB}$ 

$$\left(\frac{\nabla^2 u_A}{K} - \frac{\nabla_A K}{K} + u^A K_{AB} - u^B \nabla_B u^A\right) \mathcal{P}_C^A = \mathcal{O}\left(\frac{1}{D}\right), \quad \mathcal{P}^{AB} = \delta^{AB} + u^A u^B$$

for "flat" metric  $\eta_{AB}$  (i.e. without any horizon) [Bhattacharyya'16]

$$g_{AB} = \eta_{AB} + rac{(\emph{n}^{A} - \emph{u}^{A})(\emph{n}^{B} - \emph{u}^{B})}{\psi^{D-3}} + \mathcal{O}\left(rac{1}{\emph{D}}
ight), \quad \psiig|_{\mathsf{horizon}} = 1 ext{ is a scalar function}$$

$$S_{\text{grav.}} = \underbrace{\int_{\mathcal{M}} R \, d\text{Vol}_D}_{\text{EH without BH }} + \underbrace{\int_{\partial \mathcal{M}} K \, d\text{Vol}_{D-1}}_{\text{GHY with PT}}$$

 $\boxed{1} \Delta F_{\mathsf{grav}} \propto \Delta S_{\mathsf{GHY}}; \quad \boxed{2} S_{\mathsf{matter}}[g_{\mathsf{AB}}]; \quad \boxed{3} \mathsf{S.T.} \ \mathsf{with} \ D8/\overline{D8}/D4.$ 

## **Effective CFT within large D limit**

$$\underbrace{\int \frac{d^{d+1}x\sqrt{-g}}{16\pi \mathsf{G}_{d+1}} \left(\mathsf{R} + \frac{d(d-1)}{\mathsf{L}^2} - \frac{\mathsf{L}^2}{g_{\mathsf{S}}^2} \mathsf{F}^2\right)}_{\text{Einstein-Maxwell RN-AdS}_{d+1}} \rightarrow \underbrace{\int \frac{d^2x\sqrt{-h}\mathrm{e}^{-2\psi}}{16\pi \mathsf{G}_2} \left(\mathsf{R}_2 - \frac{1}{g_{\mathsf{S}}^2} \mathsf{F}^2 + 4(\nabla \psi)^2 + \frac{1}{\tilde{\mathsf{L}}^2}\right)}_{\text{Einstein-Maxwell-dilaton RN-AdS}_2}$$

$$\begin{split} T^{ab}\big|_{\partial \mathsf{RN-AdS}_2} &= \frac{2}{\sqrt{-h}} \frac{\delta \mathsf{S}\big|_{\partial \mathsf{RN-AdS}_2}}{\delta h_{ab}}; \quad \text{asymptotic diffeomorphism} x^\mu \to x^\mu + \epsilon^\mu \\ & (\delta_\epsilon + \delta_\Lambda) T_{tt} = \underset{\mathsf{transformation}}{\mathsf{usual}} + \underset{\mathsf{term}}{\mathsf{anomalous}}, \quad A_\mu \to A_\mu + \partial_\mu \Lambda \end{split}$$

$$\begin{array}{ccc} \text{anomalous} & \Rightarrow & c = \frac{3}{2\pi \mathsf{G}_2} = \frac{3\Omega_{d-1}}{2\pi \mathsf{G}_{d+1}} r_0^{d-1} & \Rightarrow & s_{\mathsf{CFT}} = \frac{\pi^3}{3} \mathsf{T} c = \frac{1}{4\mathsf{G}_2} & \left( \begin{array}{c} \mathsf{by} \, \mathsf{Cardy} \\ \mathsf{formula} \end{array} \right) \end{array}$$

CFT<sub>d=1</sub> reproduces   
Bekenstein-Hawking for RN-AdS<sub>d+1</sub> 
$$s = \frac{\pi^3}{3} Tc = \frac{1}{4G_2} = \frac{\text{Area}}{G_{d+1}}$$
 [Guo'16]

Sector of gravity

# Hawking-Page PT within large D limit eff. CFT

#### \*something that looks like a conclusion\*

$$F_{\text{thermal AdS}} = -pV \xrightarrow{\text{Hawking-Page}} F_{\text{BH}} = M - TS - \mu Q - pV$$

- 1.  $S_{\phi + \text{grav}} + S_{\text{gauge}}$  asymptotic RN-AdS<sub>d+1</sub> (unknown) solutions (or "near-AdS×any-compact" for top-down approach),
- 2.  $\xrightarrow{\text{large D}}$  RN-AdS<sub>2</sub> (RN-AdS<sub>3</sub>) solutions in leading 1/d order,
- 3.  $\xrightarrow{\text{border}}$  central charge of CFT<sub>1</sub> (CFT<sub>2</sub>),
- 4.  $\xrightarrow{\text{Cardy's}}$  free energy and Hawking-Page phase transition (in RN-AdS<sub>d+1</sub>),
- 5. AdS/CFT confinement/deconfinement-like PT for CH model (as it works for AdS/QCD).

#### Thank you for your attention!



Phys. Rev. D 108, 115011 (matter sector only)

System of units:  $1 = c = \hbar = k_B = \pi = e = i = -1 = 2 \text{kg}$ 

# **Holographic model**

 $\mathcal{L}_{\mathsf{CH}}$  – strongly coupled  $\Rightarrow$  consider  $\mathsf{N} \gg 1$   $\Rightarrow$   $\mathcal{Z}_{\mathsf{CH}}[J] = \mathcal{Z}_{\mathsf{AdS}}[J]$ 

The dual theory:  $\mathcal{Z}_{\mathsf{AdS}}[J] \sim \exp\left(-S_{\mathsf{AdS}}[J]\right)$  is weakly coupled  $\Rightarrow$  quasiclassical limit

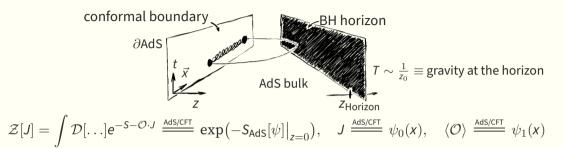
The asymptotic behavior near the conformal border ∂AdS of the dual theory fields defines the sources of the CH operators (i.e. the correlator functions)

$$X_{IJ} \stackrel{\mathsf{z} \to 0}{\longleftarrow} \frac{\sqrt{N}}{2\pi} J_{IJ} \mathsf{z} + \frac{2\pi}{\sqrt{N}} \Sigma_{IJ} \mathsf{z}^3 + \dots \quad X_{IJ} : \mathsf{AdS}_5 \stackrel{\mathsf{dual}}{\Longleftrightarrow} \Sigma_{IJ} : \mathbb{R}^{1,3}$$

Holography is the **duality** between **strongly coupled** theory on the border and **weakly coupled** (quasiclassical) bulk theory.

$$F = -T \log \mathcal{Z}_{\mathsf{CH}} \sim T \mathsf{S}_{\mathsf{AdS}} \propto \mathsf{Vol}_4 \cdot \mathcal{F} \quad \text{In homogeneous case } (\chi = \chi(\mathsf{z})): \mathcal{F} \propto \mathsf{V}_{\mathsf{eff}}[\chi]$$

## Finite temperature



EoM solutions 
$$\frac{\partial AdS}{\partial AdS}$$

boundary part  $S_{\partial AdS}$ 

quasiclassical

CFT generating function

Backup slides 14/12

# **Action of the holographic model**

$$S_{
m tot} = S_{
m grav+\phi} + S_{
m X} + S_{
m A} + S_{
m SM} + S_{
m int}, \quad S_{
m A} = -rac{1}{g_{
m z}^2} \int d^5 x \sqrt{|g|} e^\phi g^{ac} g^{bd} F_{ab} F_{cd}$$

$$S_{\mathrm{grav}+\phi} = rac{1}{I_{\mathrm{s}}^{2}} \int d^{5}x \sqrt{|g|} \mathrm{e}^{2\phi} \Big[ -R + 2|\Lambda| - 4g^{ab}\partial_{a}\phi\partial_{b}\phi - V_{\phi}(\phi) \Big], \quad a,b = 0, \dots 4$$

$$S_{
m int} = \epsilon^4 \int_{z=\epsilon} d^4 x \sqrt{|g^{(4)}|} \Biggl[ c_Y B_\mu \operatorname{Tr} \left( T_Y A^\mu 
ight) + c_W W_{k,\mu} \operatorname{Tr} \left( T_k A^\mu 
ight) + \mathcal{L}_\psi \Biggr]$$

$$S_{\text{int}} = \epsilon^4 \int_{z=\epsilon} d^4 x \sqrt{|g^{(4)}|} \left[ c_Y B_\mu \operatorname{Tr} \left( T_Y A^\mu \right) + c_W W_{k,\mu} \operatorname{Tr} \left( T_k A^\mu \right) + \mathcal{L}_\psi \right]$$

$$S_{X} = \frac{1}{k_{s}} \int d^{5}x \sqrt{|g|} e^{\phi} \left[ \frac{1}{2} g^{ab} \operatorname{Tr} \left( \nabla_{a} X^{T} \nabla_{b} X \right) - V_{X}(X) \right], \quad \nabla_{a} X = \partial_{a} X + [A_{a}, X], \quad A_{a} = 0$$

$$V_{X}(X) = \operatorname{Tr} \left( -\frac{3}{2L^{2}} X^{T} X - \frac{\alpha}{4} (X^{T} X)^{2} + L^{2} \frac{\beta}{6} (X^{T} X)^{3} + O(X^{8}) \right)$$

$$L \cdot X_{IJ} \sim \frac{\sqrt{N}}{2\pi} J_{IJ} \tilde{z} + \frac{2\pi}{\sqrt{N}} \Sigma_{IJ} \tilde{z}^3 + \dots$$

Backup slides

# Geometry

$$S_{\text{grav}+\phi} = \frac{1}{l_p^3} \int d^5 x \sqrt{|g|} e^{2\phi} \Big[ -R + 2|\Lambda| - 4g^{ab} \partial_a \phi \partial_b \phi - V_\phi(\phi) \Big], \quad a, b = 0, \dots 4$$
$$ds^2 = \frac{L^2}{\tilde{z}^2} A(\tilde{z})^2 \Big( f(\tilde{z}) d\tau^2 + \frac{d\tilde{z}^2}{f(\tilde{z})} + d\vec{x}^2 \Big), \quad \phi = \phi(\tilde{z})$$

$$f=1-rac{ ilde{z}^4}{ ilde{z}^4},\quad \phi= ilde{\phi}_2 ilde{z}^2,\quad extstyle z_{H}=rac{1}{\pi T}.$$

Backup slides 16/12

## **Effective field theory**

$$\mathcal{Z}[J] = \int \mathcal{D}\phi \, \mathrm{e}^{-S - J \cdot \phi} =: \mathrm{e}^{W[J]}$$

$$\Gamma[\langle \phi \rangle] = W[J] - \frac{\delta W[J]}{\delta J} \cdot J = \int_X d^d x \Big( \underbrace{\mathcal{K}_{\text{eff}}[\partial \langle \phi \rangle]}_{=0 \text{ if } \langle \phi \rangle = \text{const}} + V_{\text{eff}}[\langle \phi \rangle] \Big) \quad - \text{ effective action}$$

Effective potential: 
$$V_{\rm eff} = \frac{1}{{
m Vol}_4} \Gamma$$

Equation of motion (EoM): 
$$\frac{\delta\Gamma}{\delta\langle\phi\rangle} = J \stackrel{\langle\phi\rangle={\rm const}}{=} \frac{\delta V_{\rm eff}}{\delta\langle\phi\rangle} \stackrel{J=0}{=} 0$$
 gives extrema condition

#### **Effective potential**

Extrema condition:  $V_{\text{eff}}|_{\text{extrema}} = V_{\text{eff}}|_{J=0} \Leftrightarrow G_0$ ; AdS/CFT:  $G_0 \Leftarrow \text{boundary term of dual theory } S_{\partial \text{AdS}}$ 

$$\operatorname{Vol}_X V_{\operatorname{eff}}\big|_{\operatorname{extrema}} = G_0 = W[J=0] \ \ \frac{\operatorname{AdS/CFT}}{=} \ \ S_{\operatorname{AdS}}\big|_{\partial\operatorname{AdS}}^{\psi_0=0}$$

Extrema condition & duality:  $J = \psi_0 = 0$ ; duality:  $\langle \phi \rangle = \psi_1$ 

$$\frac{\delta V_{\rm eff}}{\delta \langle \phi \rangle} \ \frac{\rm AdS/CFT}{} \ \frac{\delta}{\delta \psi_1} \left( S[\psi] \big|_{\partial {\rm AdS}} \right) \Big|_{\psi_0 = 0} = 0 \quad \left( \begin{subarray}{c} {\rm with \, assumption} \\ \langle \phi \rangle = {\rm const} \end{subarray} \right)$$

gives vacuum expectation values:  $\left\{\langle\phi\rangle_{\min 1},\langle\phi\rangle_{\min 2},\dots\right\}$  — possible vacuums

Extrema positions and values 
$$\left\{\left(\langle\phi\rangle_{\min i}, V_{\text{eff}}[\langle\phi\rangle_{\min i}]\right)\right\} \Rightarrow$$
 phase transitions

#### **Gravitational Waves**

The spectrum of the gravitational waves can be estimated as (within the approach of **relativistic** velocity of the bubble walls  $v_W \sim 1$ )

$$\Omega_{\mathsf{GW}} h^2 = 1.67 \cdot 10^{-5} \kappa \Delta \left(\frac{\beta}{H_*}\right)^{-2} \left(\frac{\alpha}{1+\alpha}\right)^2 \left(\frac{g_*}{100}\right)^{-\frac{1}{3}}$$

Only scalar waves! Sound waves and turbulence are not included! We **estimate** only scalar waves produced during initial collisions.

$$f_0 = 1.65 \cdot 10^{-5} \text{Hz} \cdot \frac{f_*}{\beta} \frac{\beta}{H_*} \frac{T}{0.1 \text{TeV}} \left( \frac{g_*}{100} \right)^{\frac{1}{6}} \text{Hz}$$

 $(\Omega_{\sf GW} h^2, f_0)$ -curve is the estimation GW amplitude (peak value). It does not contain the spectral shape  $S(f_0)$  (in this case  $S(f_0 = f_0^{\sf peak}) = 1$ ).

## Temperature estimations

Experimental restrictions  $\Leftrightarrow$  mass of the **lightest predicted** particle.

$$\Sigma_{IJ} = \xi^{\top} \begin{bmatrix} \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \varsigma \end{pmatrix} + \eta_i \tilde{T}_i \end{bmatrix} \xi \quad \frac{\mathsf{AdS/CFT}}{\mathsf{dual to}} \quad \mathsf{X}_{IJ} \to \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & \chi \end{pmatrix}$$

 $m_n \sim m_{\delta \chi}$  fluctuation mass  $\sim$  slope of the "hat".

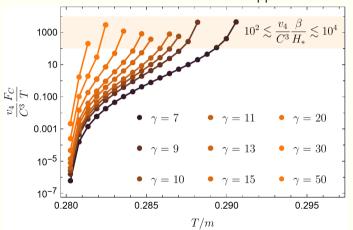
$$\chi(\mathbf{z}) \to \chi(\mathbf{z}) + \delta \chi(t, \vec{\mathbf{x}}, \mathbf{z}) \quad \Rightarrow \quad \mathsf{EoM}_{\mathbf{z}}[\chi] \to \mathsf{EoM}_{t, \vec{\mathbf{x}}, \mathbf{z}}[\chi + \delta \chi] \qquad \xrightarrow{\partial \mathsf{AdS}} \qquad \boxed{2m^2 = \phi_2}$$

$$T = rac{1}{\pi} rac{1}{z_{\mathsf{H}}} \quad \Rightarrow \quad T = rac{m}{\pi} \sqrt{rac{2}{\phi_2}}, \quad z_{\mathsf{H}}^2 = rac{\phi_2}{2m^2}$$

# **Bubble free energy**

Free energy of a bubble:

 $F[V_{
m eff}] = \frac{
m tnin \, walls}{
m approximation} \, 4\pi R^2 \mu - rac{3\pi}{4} R^3 \left( \mathcal{F}_{
m out} - \mathcal{F}_{
m in} 
ight)$ 



 $F_{\rm C} \stackrel{\rm def}{=} F(R_{\rm C})$ : if  $R > R_{\rm C}$ , bubbles grow and **PT occurs**.

$$\frac{\beta}{H*} \sim \frac{F_C}{T} + \mathcal{O}(T)$$

 $1/\beta \sim$  appear  $\rightarrow$  collide time  $1/H_* \sim$  universe expansion

$$v_4 \sim 10^{-1} \ll 1, \quad c \sim 1$$

$$10^3 \gtrsim \frac{\beta}{H_*} \gtrsim 10^5$$

## Holographic effective potential

$$\mathcal{Z}_{\mathrm{CH}}[J] = \int \mathcal{D}\varphi \, \exp\left(-\mathsf{S}[\varphi] - \int d^4x \, \varphi(x) \, J(x)\right) \stackrel{\mathsf{def}}{=} e^{-W[J]}$$

$$\langle \varphi \rangle = \frac{\delta W[J]}{\delta J} \bigg|_{t=0}$$
,  $\Gamma[\langle \varphi \rangle] = W[J] - \int d^4x \, \frac{\delta W[J]}{\delta J(x)} J(x)$  – Effective Action

extrema condition

$$\mathcal{Z}[J] \stackrel{\underline{\mathrm{AdS/CFT}}}{=\!=\!=\!=\!=} \mathcal{Z}_{\mathrm{AdS}} \stackrel{\mathrm{quasiclassical}}{=\!=\!=\!=\!=} e^{-S_{\mathrm{AdS}}}|_{\partial \mathrm{AdS}} - \frac{\mathrm{quasiclassical}}{\mathrm{non-perturbative}}$$

$$\left| V_{
m eff} = -rac{1}{{
m Vol}_4} {
m S}_{
m AdS} 
ight|_{\partial {
m AdS}} - rac{{
m boundary \, term \, of \, the \, bulk \, theory \, defines \, quantum \, effective \, potential}$$

#### "Extrema" curves

$$\frac{\delta S_{\chi}}{\delta \chi} = 0 \Rightarrow \chi \xrightarrow{z \to 0} Jz + \left(\sigma - \left(\frac{3}{2}J^3 + \phi_2 J\right)\log z\right)z^3 + o(z^5) - \frac{\text{give the sourses}}{\text{for CFT operators}}$$

Knowing the *extrema* of the effective potential and its *values* at these points, we can judge abut the phase transition

$$V_{\rm eff} = -\frac{1}{{\rm Vol}_4} S_\chi \Big|_{\partial {\rm AdS}} \Rightarrow \begin{array}{l} {\rm from \ EoM} \\ {\rm for \ effective} : {\rm Vol}_4 \frac{\delta V_{\rm eff}}{\delta \langle \varphi \rangle} = J \Rightarrow \begin{array}{l} {\rm extrema\ condition\ is} \\ {\rm absence\ of\ sources} \end{array} \Rightarrow J = 0$$

$$\overbrace{\chi \xrightarrow{z \to 0} \sigma z^3 + o(z^5)}^{\text{"extreme" solutions}} \text{ must give } \overbrace{\frac{\delta V_{\text{eff}}}{\delta \langle \varphi \rangle}}^{\text{extrema}} = 0 \Rightarrow \text{ a new condition for } \phi_2 \text{ and } \langle \varphi \rangle$$

$$T \sim \frac{1}{\sqrt{\sigma_2}}, \quad \frac{\delta V_{\text{eff}}}{\delta \langle \varphi \rangle} = 0 = \frac{\delta}{\delta \sigma} S_{\chi} \left[ \chi_{\text{Sol.}}(z; J, \sigma) \right] \Big|_{I=0} \quad \Rightarrow \quad \{\sigma_1, \dots, \sigma_n\} - \text{extrema}$$

Effective theory approach to  $\langle \varphi \rangle$ , vacuum average of the effective theory

#### **Nucleation ratio**

Baryogenesis generates enough asymmetry (enough efficient) if there is one bubble per Hubble volume

Nucleation: 
$$AT^4e^{-\frac{F_c}{T}}$$
  $\sim H^4(T) = \left(\frac{T^2}{M_{Pl}}\right)^4$  - Expansion of the Univerce Bubbles produced per time×space volume  $1/(\text{Hubble time} \times \text{volume})$ 

 $F = F[\langle \varphi \rangle, R]$  – Free energy of the bubble; R is the radius of the bubble

Hubble horizon (time, volume, radius) — speed of receding object behind it is greater than the speed of light (Don't confuse with cosmological horizon)

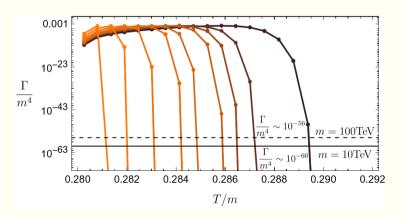
Bubble appears with a certain size. It defines with "micro-physics". If its radius is grater, then critical one  $\frac{\partial F}{\partial R}\Big|_{R_c \stackrel{\text{def}}{=} R}$ , the bubble grow. Otherwise, it bursts.

It gives  $F_C \stackrel{\text{def}}{=} F(R_C)$  and defines nucleation ratio and "viability of the model".

#### **Estimations of the nucleation ratio**

$$\Gamma \sim H_*^4, \quad \frac{\Gamma}{m^4} \sim \frac{H_*^4}{m^4} \propto \frac{m^4}{M_{
m Pl}};$$

 $F_C$  is defined with an error, so  $e^{\frac{F_C}{T}}$  has large error



# Potentials of CFT and the dual theory

$$V_\chi = a_2 \chi^2 + a_4 \chi^4 + a_6 \chi^6, \quad a_2 < 0, \, a_4 < 0, \, a_6 > 0 \quad \text{no barrier}$$
  $V_{ ext{eff}} = b_2 \langle \varphi \rangle^2 + b_4 \langle \varphi \rangle^4 + b_6 \langle \varphi \rangle^6, \quad b_2 > 0, \, b_4 < 0, \, b_6 > 0 \quad ext{there's a barrier}$ 

in details:

- $V_{
  m eff} = V_{
  m eff}[\langle arphi 
  angle]$  describes a quantum objects at the border.  $V_\chi$  is a dual classical potential in the bulk.
- $V_{
  m eff} = -rac{1}{{
  m Vol}_4} S_{
  m AdS} \Big|_{\partial {
  m AdS}}$  includes the solutions of the EoM  $rac{\delta S_\chi}{\delta \chi = 0}$  in bulk. In other words,  $V_{
  m eff}$  includes physics of AdS

# **Conditions for the dual theory potential**

$$V_\chi(\chi)=rac{m^2}{2}\chi^2-rac{D}{4L^2}\lambda\chi^4+rac{\lambda^2\gamma}{6L^2}\chi^6$$
 is the expantion of a more general theory

#### Suggestions:

- The potential  $V_{\chi}$  always has true vacuum with  $E_{\min}$  ( $V_{\chi} \xrightarrow{\chi \to \pm \infty} \infty$ ). So we may use any even power  $\chi^n$  instead of the last term  $\chi^6$ .
  - The expansion of  $V_{\chi}$  has certain sign of the second term  $\lambda > 0$  (the first one  $m^2$  chosen for the theory to be conformal in AdS).
- Higher orders of the expansion don't give new minima at the considered temperatures.

The certain parametrization has been chosen with respect to the "symmetries"

"Scale invariace", 
$$L \to L'$$
 Conformality near the AdS border the coefficents  $\chi \to \sqrt{\lambda} \chi$ ; the AdS border  $\Delta_+ = 1$   $\Delta_- = 1$   $\Delta_+ = 3$   $\Delta_+ = 3$ 

D is for the Large D limit. But its usage doesn't give any results.

(to keep interaction constants finite at  $D \to \infty$ ) Effective theory approach

#### **SM - CH model interactions**

$$F = \frac{\text{thin walls}}{\text{approximation}} 4\pi R^2 \mu - \frac{3\pi}{4} R^3 \left( \mathcal{F}_{\text{out}} - \mathcal{F}_{\text{in}} \right) - \text{physical units are required}$$

- Fix the Parameters (Interaction with Standard Model bulk gauge fields)
- Physical Units (Infrared Regularization and finite temperature "radial" heavy fluctuations)

$$W^{lpha}_{\mu}J^{lpha\,\mu}_{
m L}+B_{\mu}J^{\mu}_{
m Y}\quad\Leftrightarrow\quad J^{\mu}\sim {\it A}^{\it M}$$
 – bulk  ${\cal G}$  gauge field

The physical values can be estimated without gauge field:

$$\Sigma_{IJ} = \langle \bar{\Psi}_I \Psi_J \rangle = \xi^{\top} \begin{bmatrix} \begin{pmatrix} 0_{4 \times 4} & 0 \\ 0 & X \end{pmatrix} + \eta_i \tilde{T}_i \end{bmatrix} \xi \quad \Leftrightarrow \quad \frac{1}{T} \propto \sqrt{\phi_2} \sim \mu_{\mathrm{IR}} \sim m_{\eta} \gtrsim 10 \,\mathrm{TeV}$$
 $m_{\eta} \leftarrow X \rightarrow X + \delta X$  – correction of the background field  $\Rightarrow \eta$  – pNG boson

#### CH gauge field

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{CH} + B_{\mu} \operatorname{Tr} \left( T_{Y} \hat{J}^{\mu} \right) + W_{k,\mu} \operatorname{Tr} \left( T_{k} \hat{J}^{\mu} \right) + \sum_{r} \bar{\psi}_{r} \mathcal{O}_{r} + \text{h.c.}$$

$$= \mathcal{L}_{interactions}$$

$$SO(5) \times U(1) : A_{M} = A_{M}^{K} T^{K} + A_{M} \times T_{Y}$$

$$SO(5) \rightarrow SO(4): \underbrace{A_{M}^{K}T^{K}}_{\in SO(5)} \rightarrow \underbrace{A_{M}^{a}T^{a}}_{\in SO(4)} + \underbrace{A_{M}^{i}T^{i}}_{\in SO(5)/SO(4)}$$
 
$$SO(4) \cong SU(2) \times SU(2): A_{M}^{K}T^{K} = \underbrace{A_{M}^{k,L}T_{L}^{k}}_{\in SU(2)_{L}} + \underbrace{A_{M}^{k,R}T_{R}^{k}}_{\in SU(2)_{R}}$$
 conserved currents:  $\hat{J}_{\mu} \stackrel{\text{dual}}{\Longleftrightarrow} A_{\mu}(t,x,z)\Big|_{z=0}^{\partial \text{AdS}}$ , holographic gauge:  $A_{z}=0$ 

$$\mathcal{O} \overset{\mathsf{dual}}{\Longleftrightarrow} \mathcal{J}(\mathsf{A},\Psi,\phi,\ldots)$$
 — composite operators of the CH fields